

# A French participation to SUMIRE/PFS

Marseille groupe LAM/CPPM

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# Science Interest

Science Interests in France

# Science interests

Strong interest in Cosmology/Extragalactic research

□ In particular

- BAO survey; target selection using photo-z
- BAO using the Ly-alpha forest of distant quasars
- Link weak-lensing with galaxy redshift surveys: halo mass measurements, cosmological constraints
- Cosmic magnification
- Strong lensing systems
- High-z galaxy search

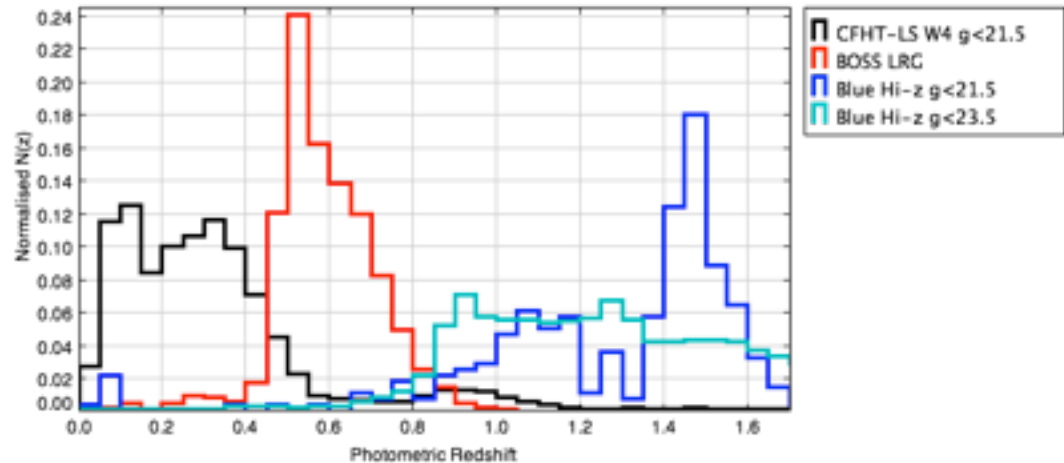
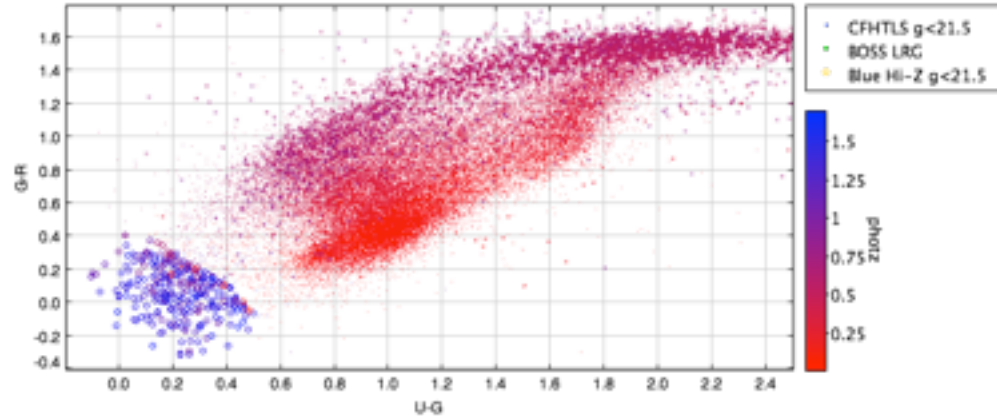
# Marseille Heritage: Hardware & Science

- VIMOS instrument (build by LAM + Italians)
- VVDS, VIPERS survey
- CFHT-LS photo-z, Cluster catalogue, SNLS spectroscopy
- COSMOS, zCOSMOS
- SNAP spectrograph/SNAP science case
- tools to prepare DE experiments (as Cosmos Mock Catalogue)
- BOSS (link WL+spectro redshift; test ELG selections)

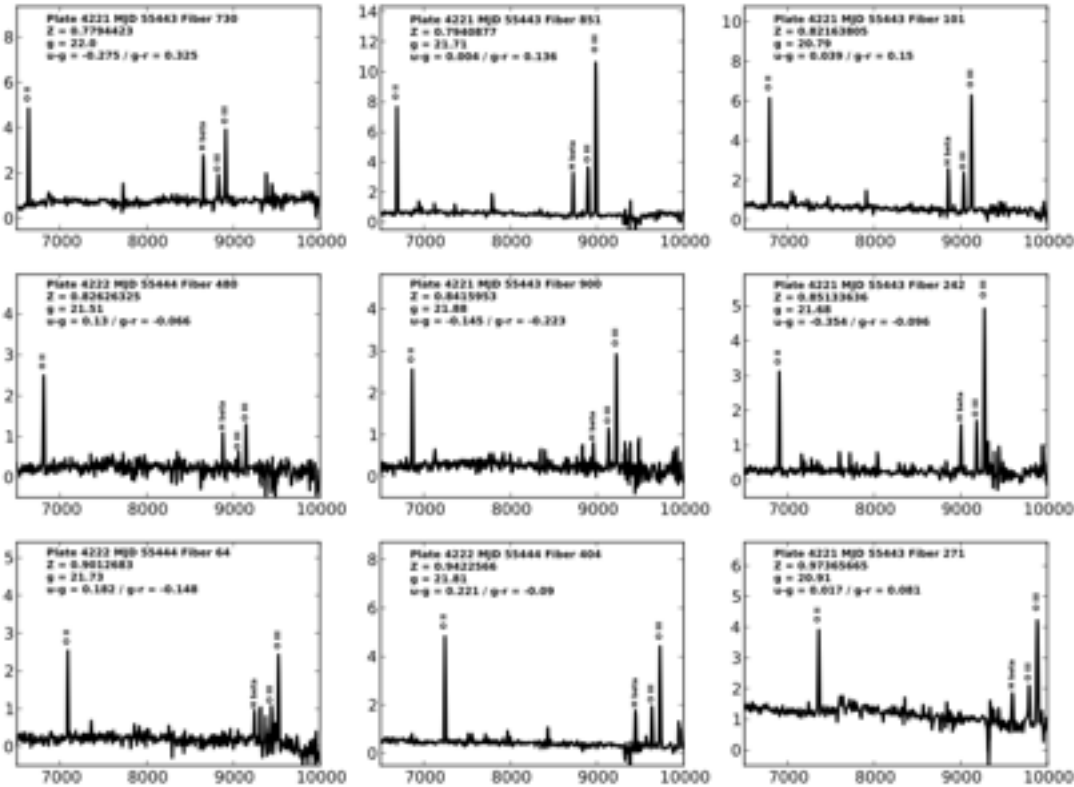
# BOSS Ancillary ELG survey

(JP Kneib, J.Companat, A.Ealet, S.Escoffier)

- ugr selection for  $0.7 < z < 1.7$  to be tested with the BOSS spectrograph this fall 2010
- other color selections possible like gri color (similar to DEEP2)



# BOSS ELG Ancillary program (2000 targets $0.7 < z < 1.7$ )

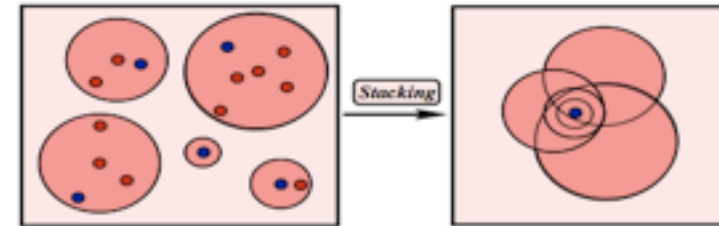
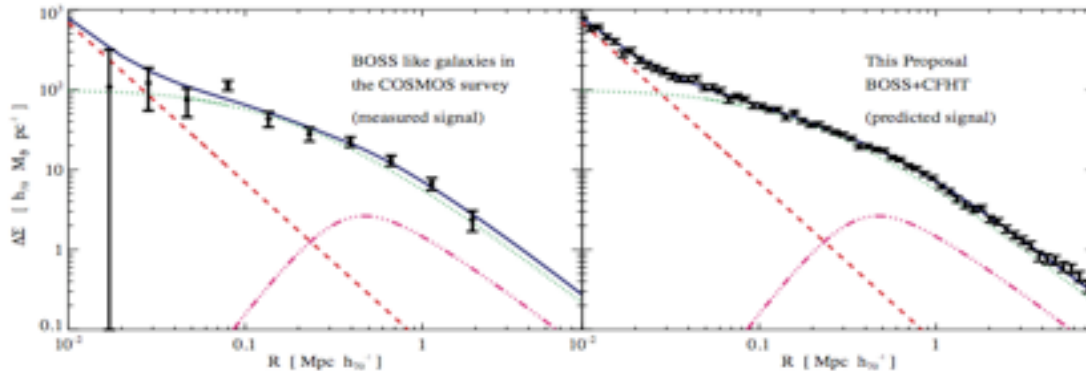


Issues of the ELG selection: quasar contamination (but it can be reduced if needed by measuring time-variability of quasars)

# WL & Spectroscopy

- Combine WL & Spectroscopic information
- Framework: Halo Model
- Can put stringent constraints on the mass distribution but also Cosmology

## Galaxy-galaxy lensing technique



CFHT WL forecast  
measurement  
in the Stripe 82  
(on-going project)

SUMIRE/PFS

Hardware contribution (Marseille)

Spectrograph



# The Marseille technical heritage

- ▣ LAM\* heritage on large survey spectroscopy = VIMOS builder ( PI VVDS : O. Lefevre director of LAM)
- ▣ An integrated technical team between LAM and CPPM\* in the context of SNAP/JDEM, to build an IFU spectrograph with slicer
- ▣ A full demonstrator has worked in 2006/2008 (see SPIE 2006/2008)
- ▣ Currently, the team has made some preliminary studies for BigBoss on a 5000 fibers spectrograph, and works on EUCLID (ESA mission)

\*LAM /INSU (Laboratoire d'Astrophysique de Marseille)

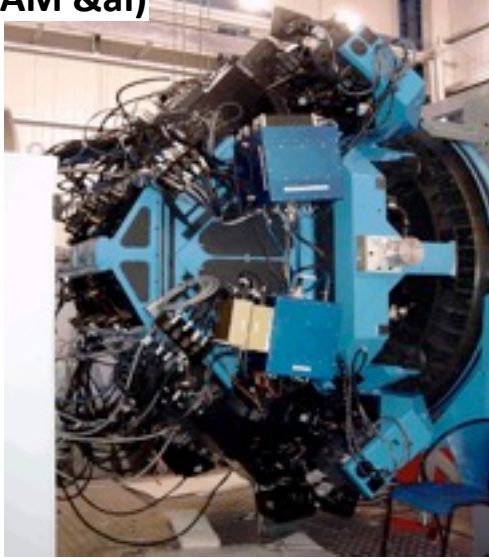
- Cosmology group head by JP Kneib

\*CPPM/IN2P3 (Centre de physique des particules de Marseille)

-Dark energy research group leaded by A.Ealet

# Marseille expertise spectroscopy/VLT-ESO

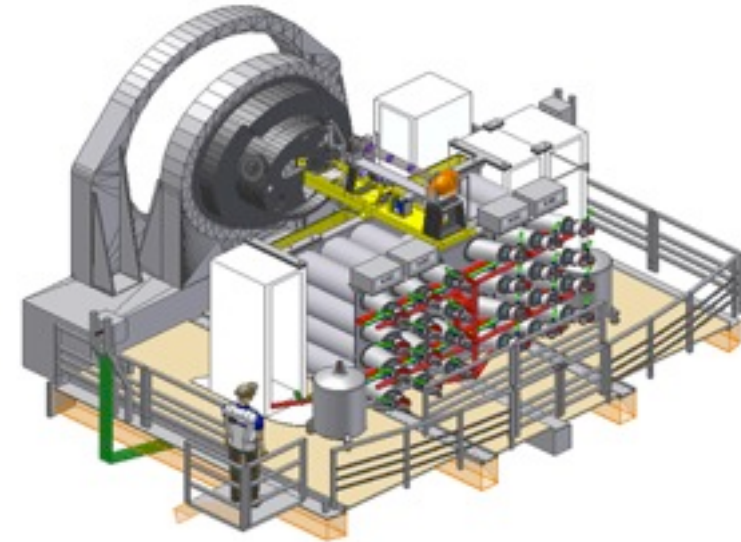
VIMOS (LAM &al)



6400 Fibers packed in 80 modules



MUSE (CRAL &al)



24 spectrographs:  
24 detectors  
No mechanism

# Deliverables

- Marseille proposes to study and to deliver
  - Spectrograph concept
  - Spectrograph optics
  - Spectrograph mechanics
  - Spectrograph cryostat
  - AIT/V
  - Software

The Marseille group has done a first technical and feasibility study for a spectrograph

# Work done in Marseille 2010

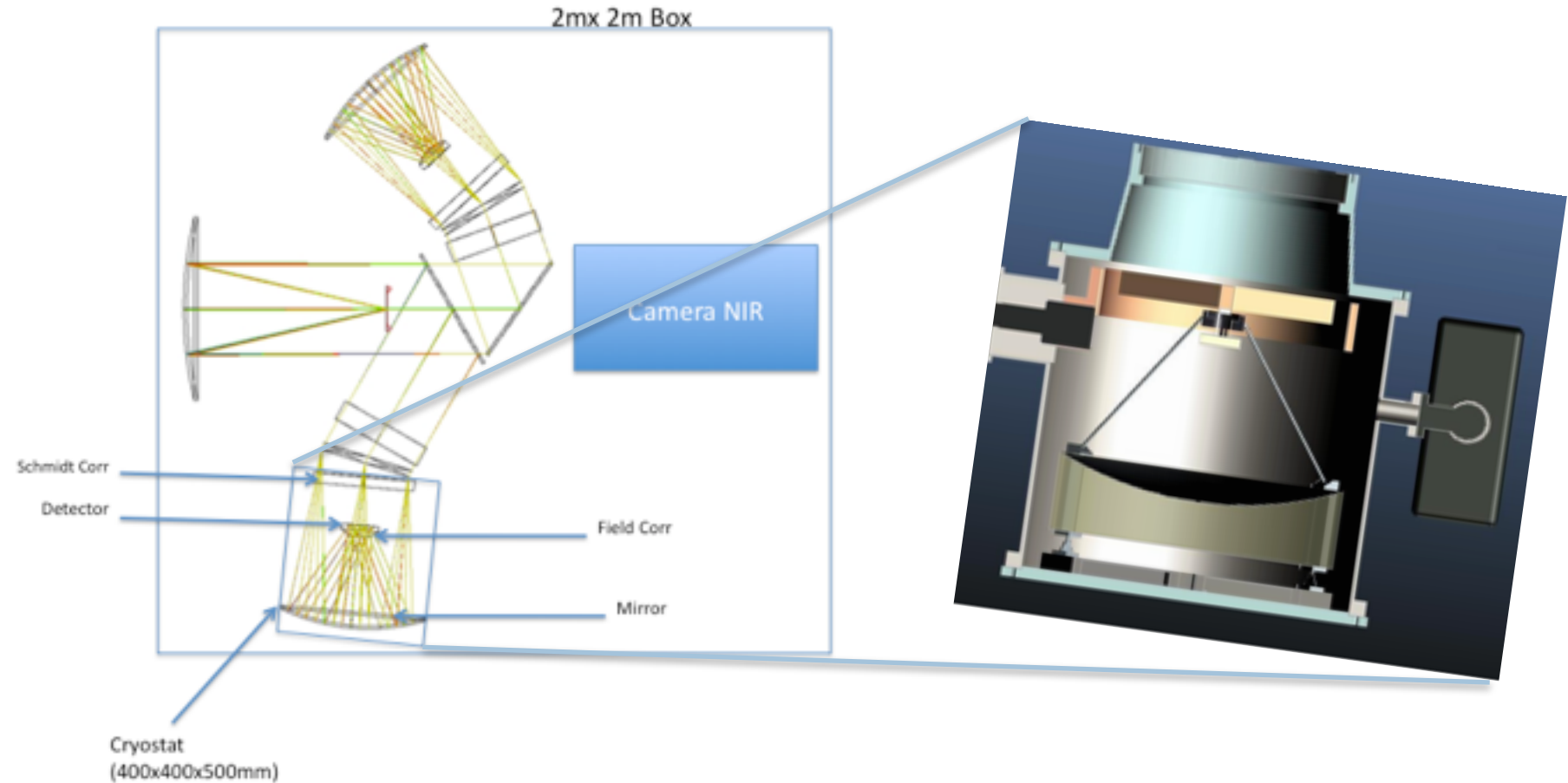
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- Optical design using Schmidt concept
- Industrial evaluation for the manufacturing
- Preliminary development plan
- Optical performance evaluation

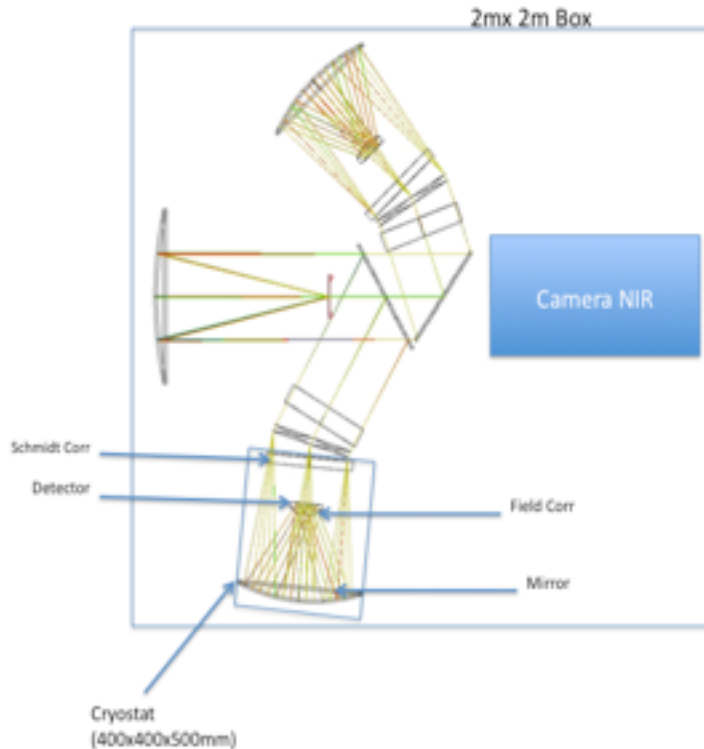
# Optical concept specification

- Wavelength band: (0.36 – 0.6 $\mu\text{m}$ ) and (0.6 – 1 $\mu\text{m}$ )
- 100 $\mu\text{m}$  fiber core
- 120mm entrance slit length
- 4kx4k detector (15 $\mu\text{m}$  pix)
- G: 0.5 (ie: 100 $\mu\text{m}$  fiber core imaged on 50 $\mu\text{m}$ : 3.33pix)
- F/2.2 collimator; F/1.1 camera
- Spectral dispersion:
  - Blue: 0.58 $\text{\AA}$ /pix
  - Red: 0.97 $\text{\AA}$ /pix

# Spectrograph Concept



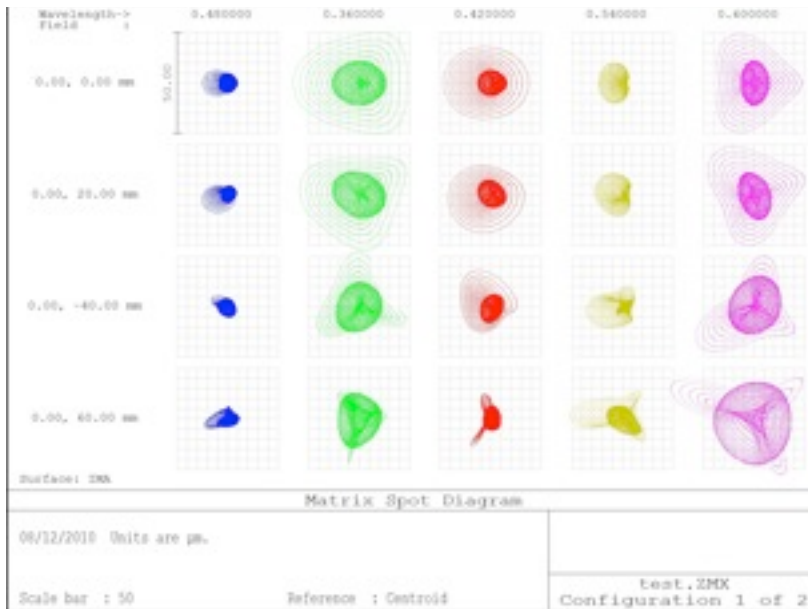
# Optical Layout



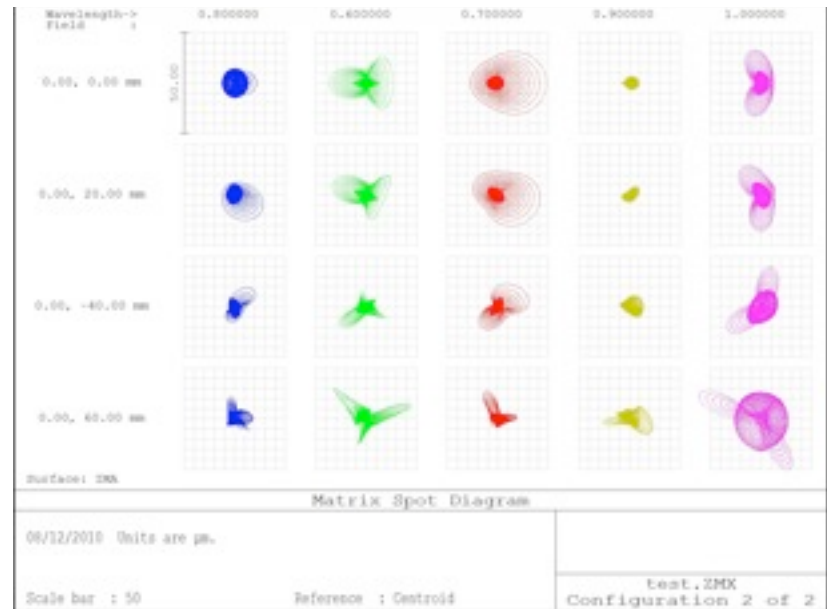
- Curved fiber slit
- Spherical collimator
- Two dichroics
- Collimator corrector dedicated to each channel (wavelength optimisation)
- Grating (280mm)
- Line density :
  - Blue: 777 ln/mm
  - Red: 472 ln/mm
- IR camera slot available
- 4 aspherical surfaces

# Spot diagram

## Blue Channel (box : fiber core size)



## Red Channel (box: fiber core size)



Red channel better than blue  
Spot smaller than fiber core image



# Throughput

## UV/visible/red

**Optics: 93 %**

**Grating: 80%**

**Detector: 90%**

**Shadowing: 85%**

**total throughput: 57 %**

## □ Included:

Specular reflexion

Internal transmittance

Dichroics

Detector support shadowing

## □ Not included:

Telescope

Fiber

Fiber couplings (FRD.....)

# Conclusion

- France is interested in participating to PFS
- Strong expertise of the Marseille group in building spectrographs
- Marseille has made a first technical study: a spectrograph concept exists and is viable
- Need to assess the number of fibers, the real specifications for the spectrograph, the interface with fibers and detectors
- Ready to participate in a definition phase