



University
of Victoria



TOHOKU
UNIVERSITY

Raven

a scientific and technical
Multi-Object Adaptive Optics (MOAO)
demonstrator

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& all RAVEN team*



Why MOAO?

- AO corrected FoV is limited by anisoplanatism

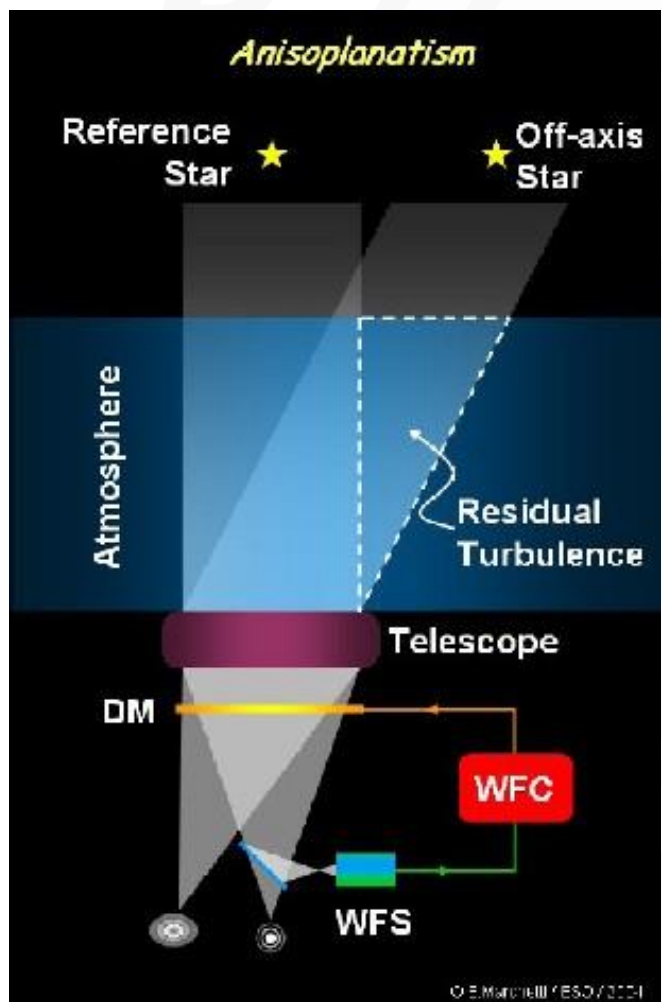
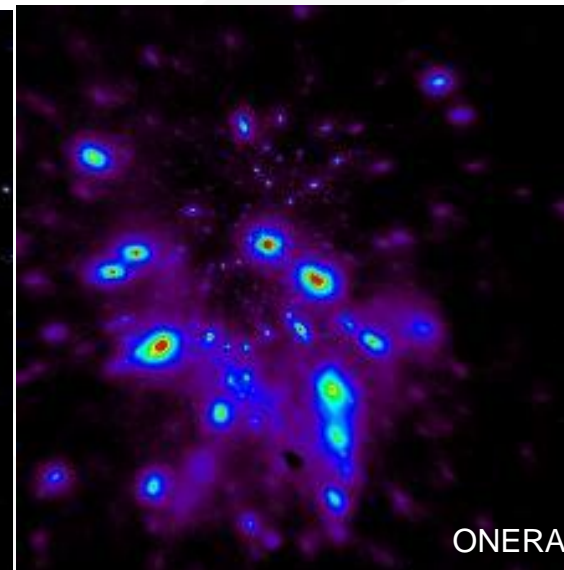
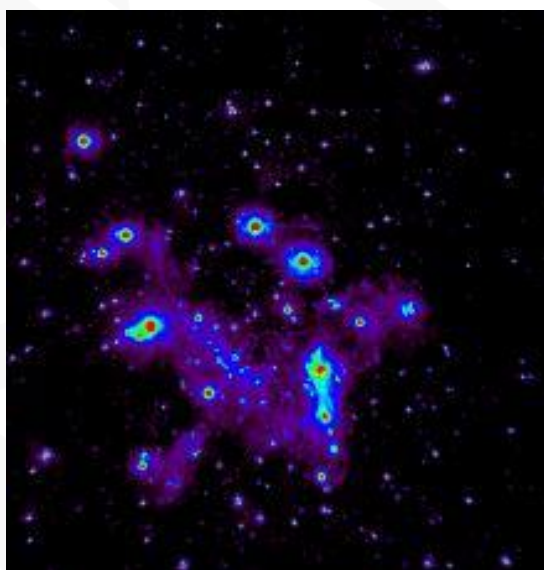


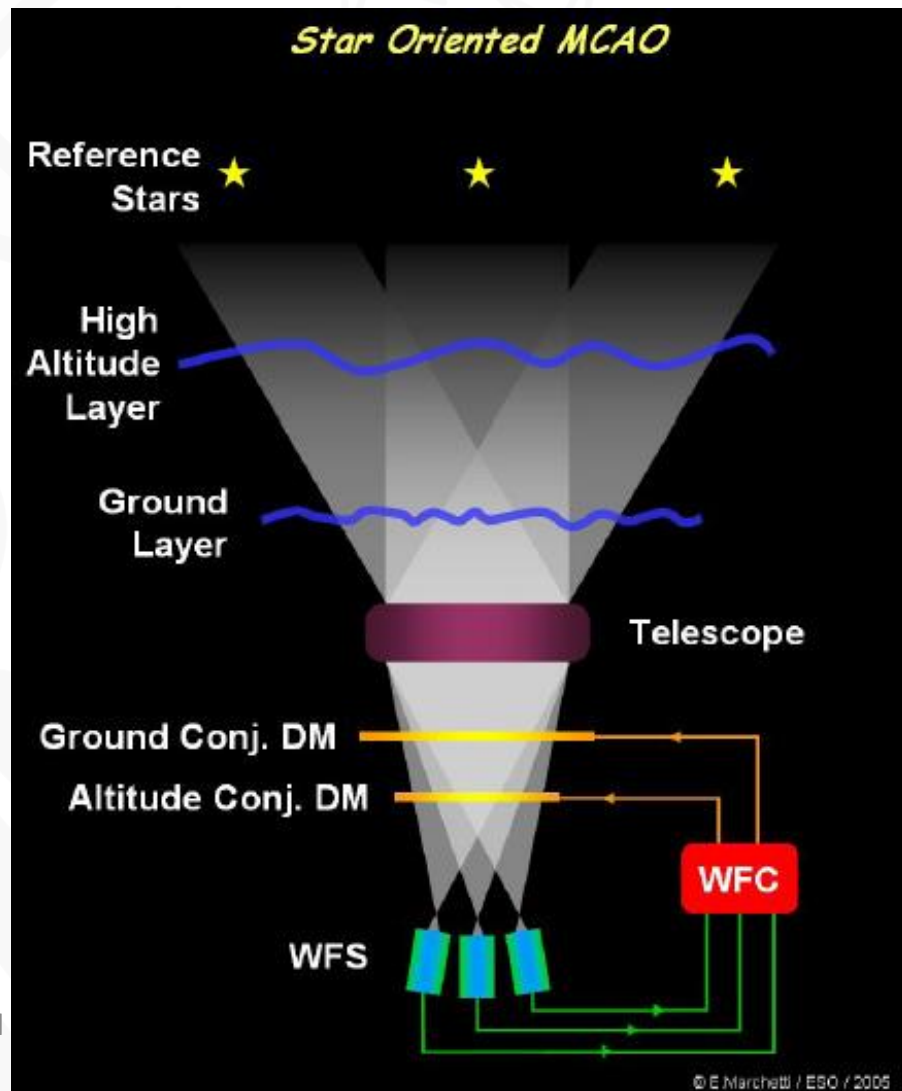
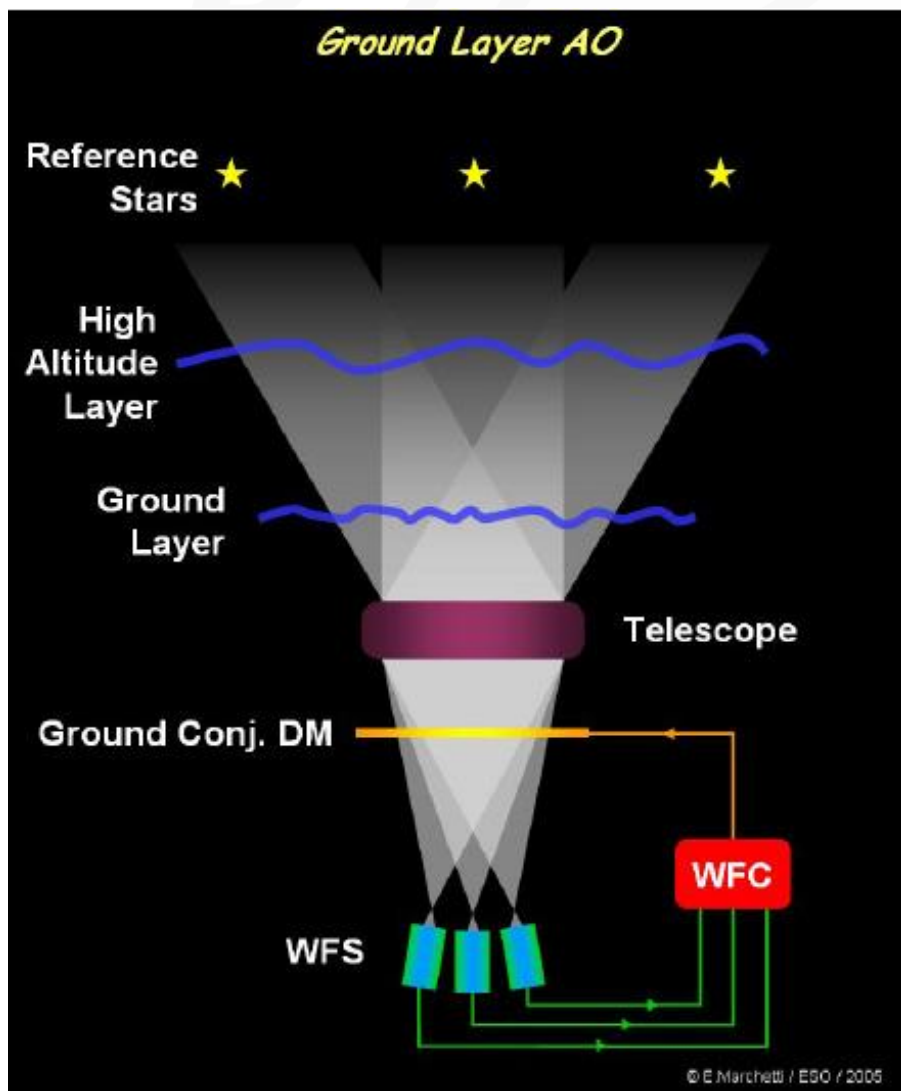
Image of galactic centre:
w/o anisoplanatism *w/ anisoplanatism*





How to increase the corrected FoV?

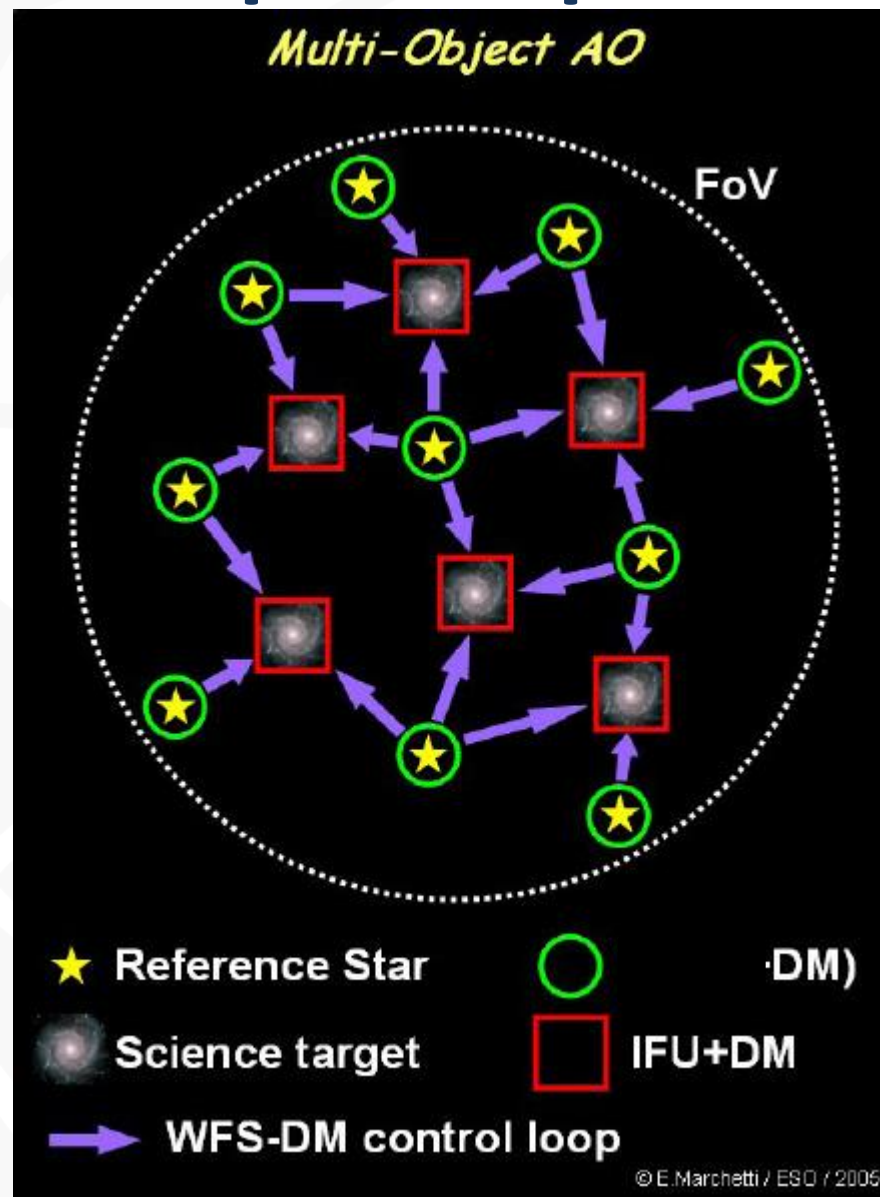
- Increase the # of guide stars and DMs





Multi-Object Adaptive Optics

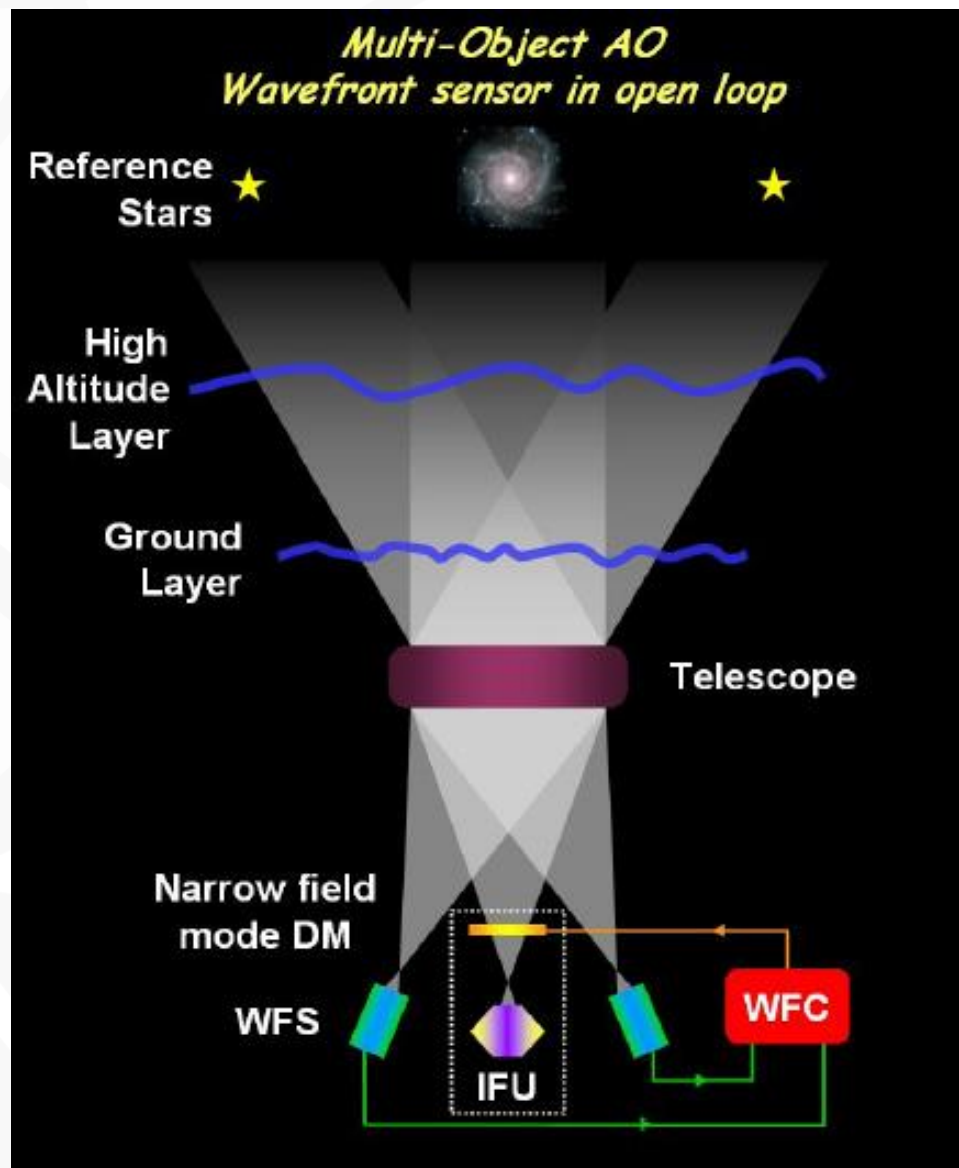
- We don't need to correct the whole FoV, just a few patches
- MOAO uses multiple WFSs to reconstruct a volume of turbulence
- A single DM can be used to apply the optimal AO correction for any single location in the field





Multi-Object Adaptive Optics

- We don't need to correct the whole FoV, just a few patches
- MOAO uses multiple WFSs to reconstruct a volume of turbulence
- A single DM can be used to apply the optimal AO correction for any single location in the field
- MOAO correction is applied in open loop

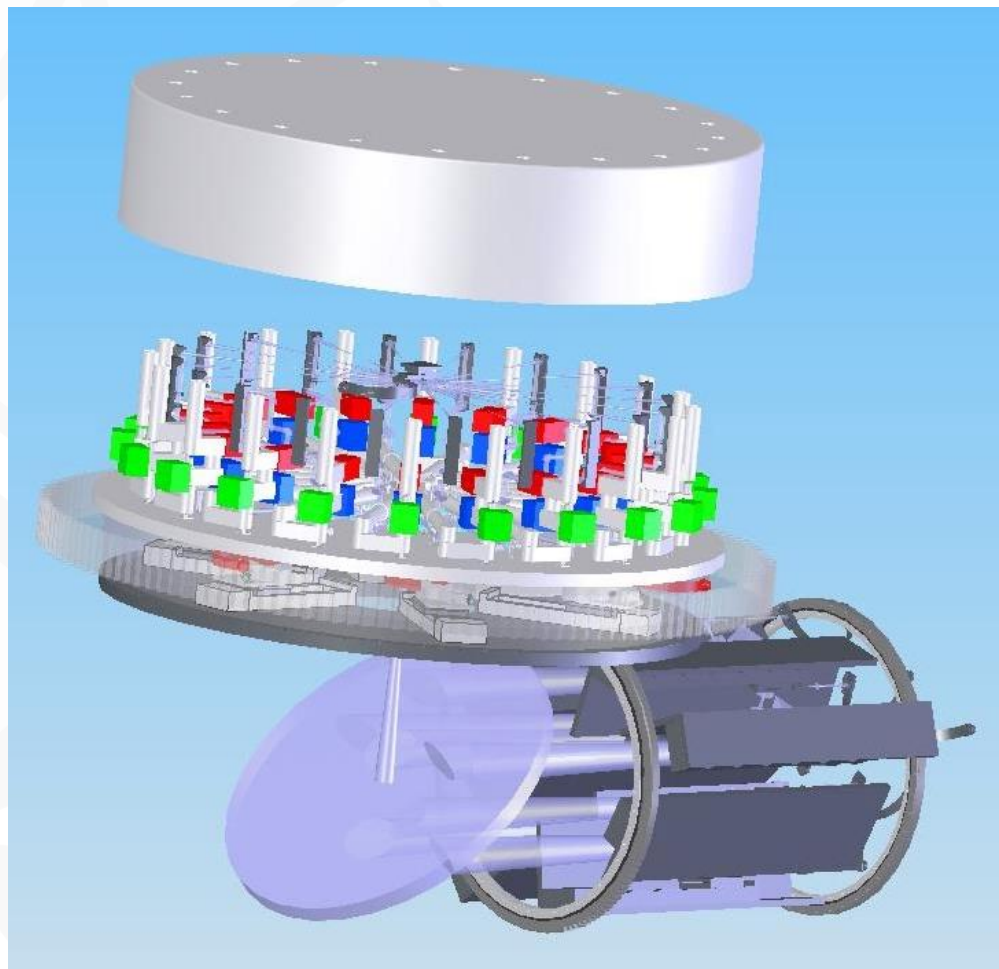




MOAO Challenges

Pickoff system

- *Up to 20 deployable pickoff mirrors are planned for ELTs (IRMOS, MOSAIC)*



IRMOS (Eikenberry & Andersen 2006)



MOAO Challenges

Pickoff system

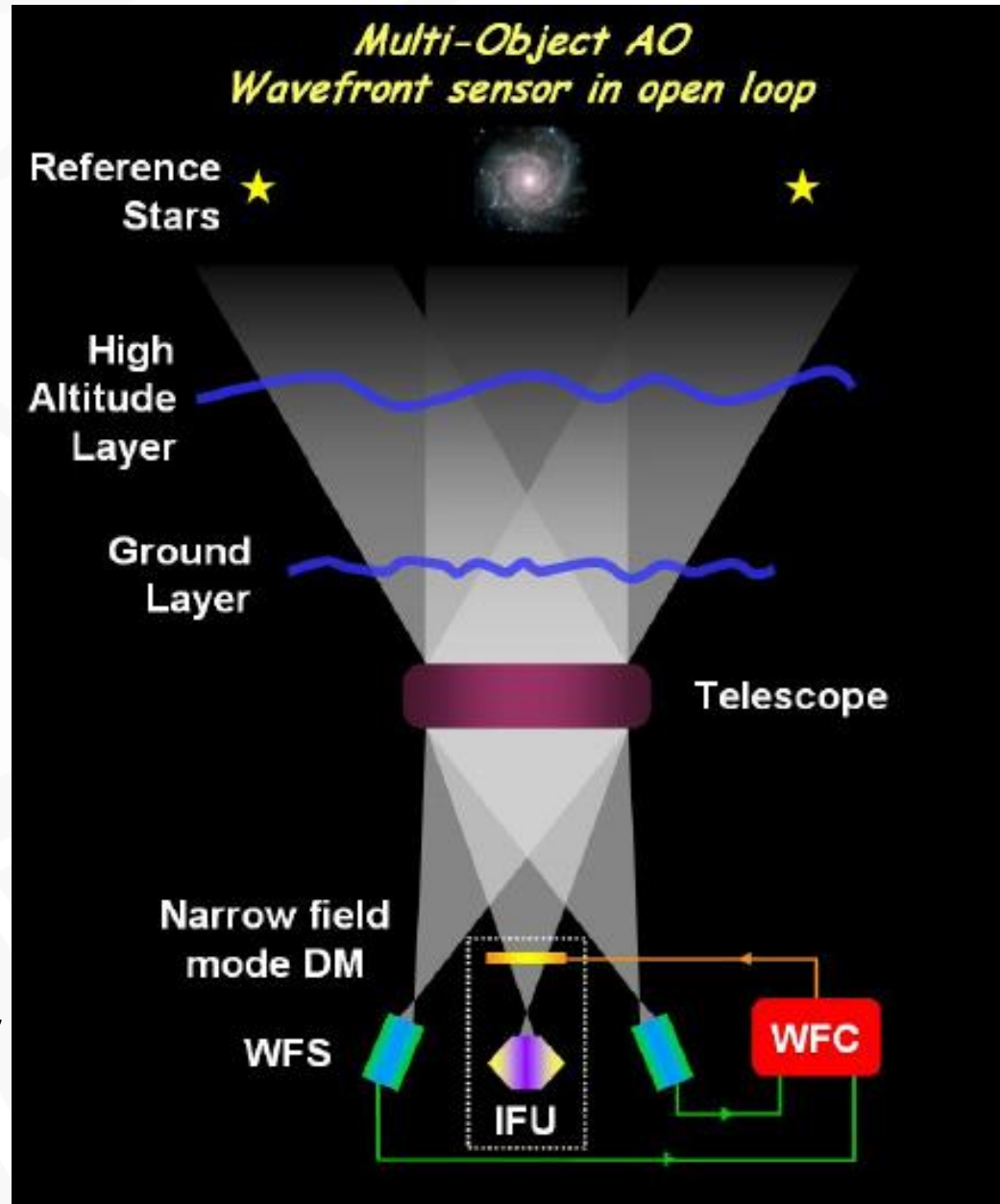
- *Up to 20 deployable pickoff mirrors are planned for ELTs (IRMOS, MOSAIC)*

Tomography

- *Finite number of WFSs*
- *Requires turbulence model*
- *Noise & static error propagation*

Open-loop control

- *Wide-field WFS*
- *DM linearity & repeatability*
- *DM/WFS calibration*





Goals of Raven

Technical & Science demonstrator

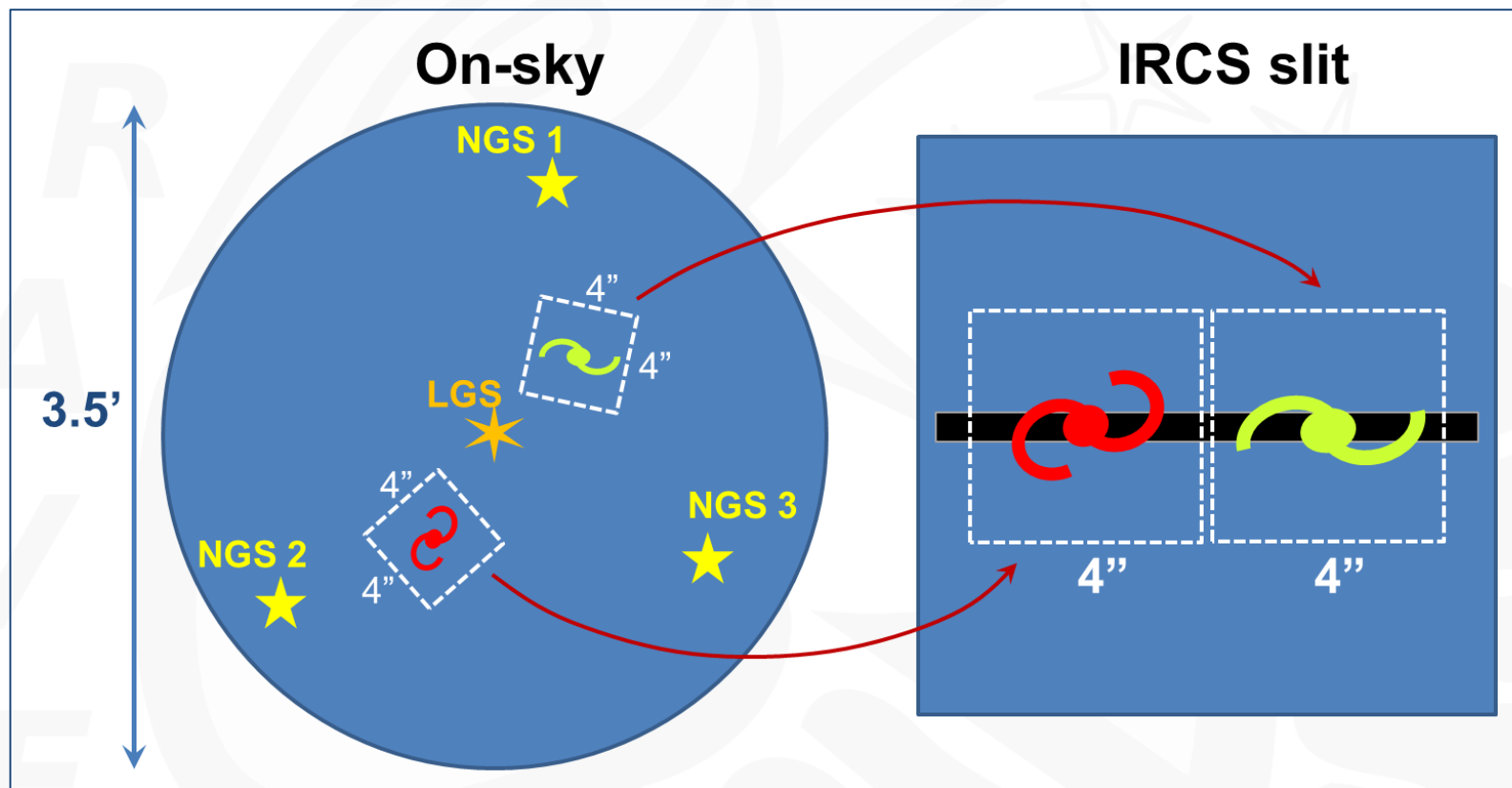
- *Prepare MOAO instrumentation for ELTs*
- ***Get first science results from MOAO***

Raven = 1st MOAO system on an 8m-class telescope feeding a NIR science instrument (IRCS):

- *Founded by CFI (Canadian Fund for Innovation) ~ 4M\$*
- *Designed and built by UVic, HIA/NRC and INO*
- *Fast track project:*
 - CoDR in Mar 2011
 - Shipped to Subaru in Jan. 2014,
 - First night in May 2014



Raven Concept

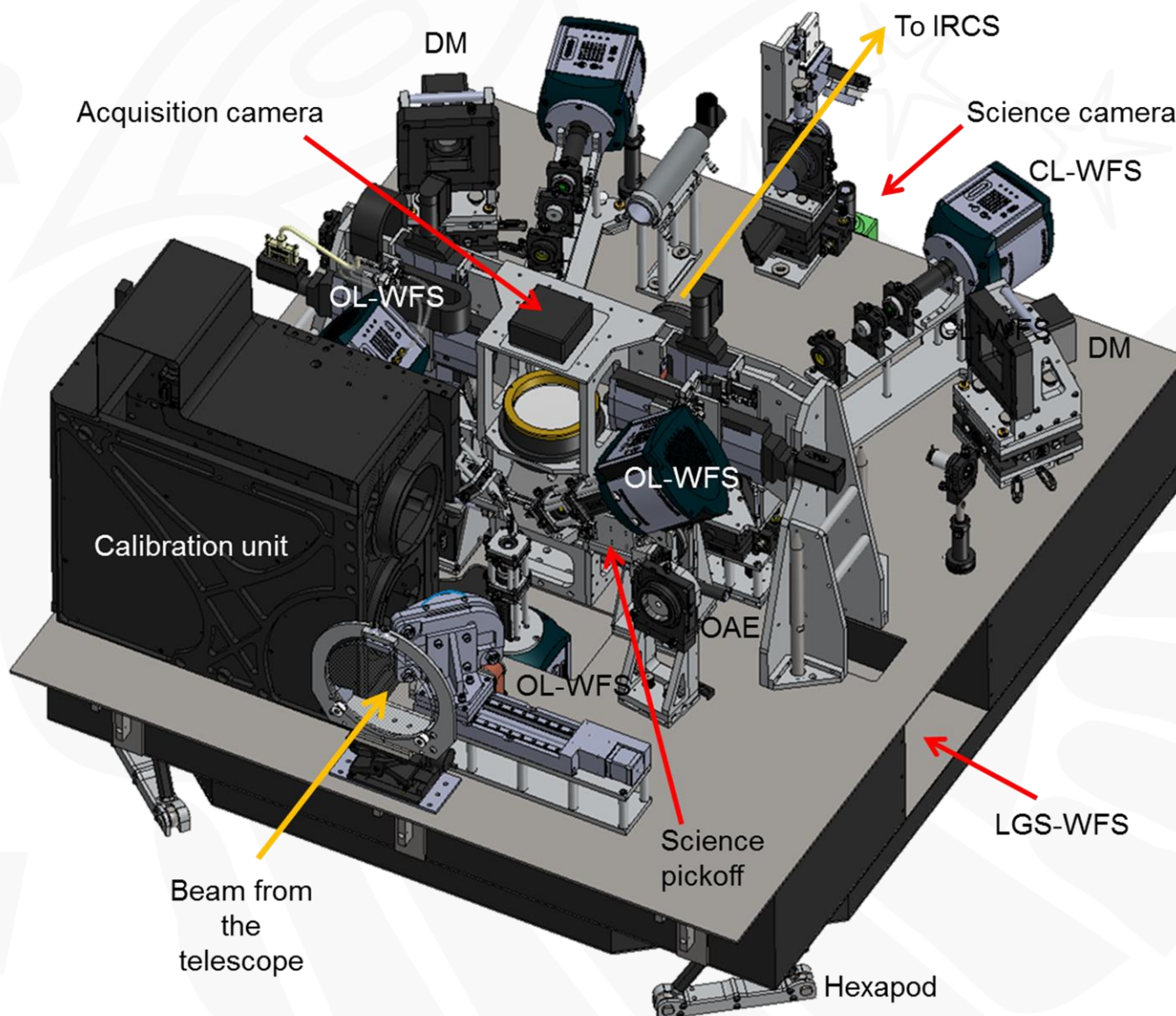


- 2 science targets are reimaged on the IRCS slit
- Each target can be positioned on the slit by moving the pickoff arms
- Each target can be rotated with the K-mirror





CAD Model





Calibration Unit = Subaru Telescope in a box!



**Source
light pipe**

Pinhole array

- 2.9'FoV
- Field rotation

LGS source

- 80-180km range

Phase screens

- 5 & 10km

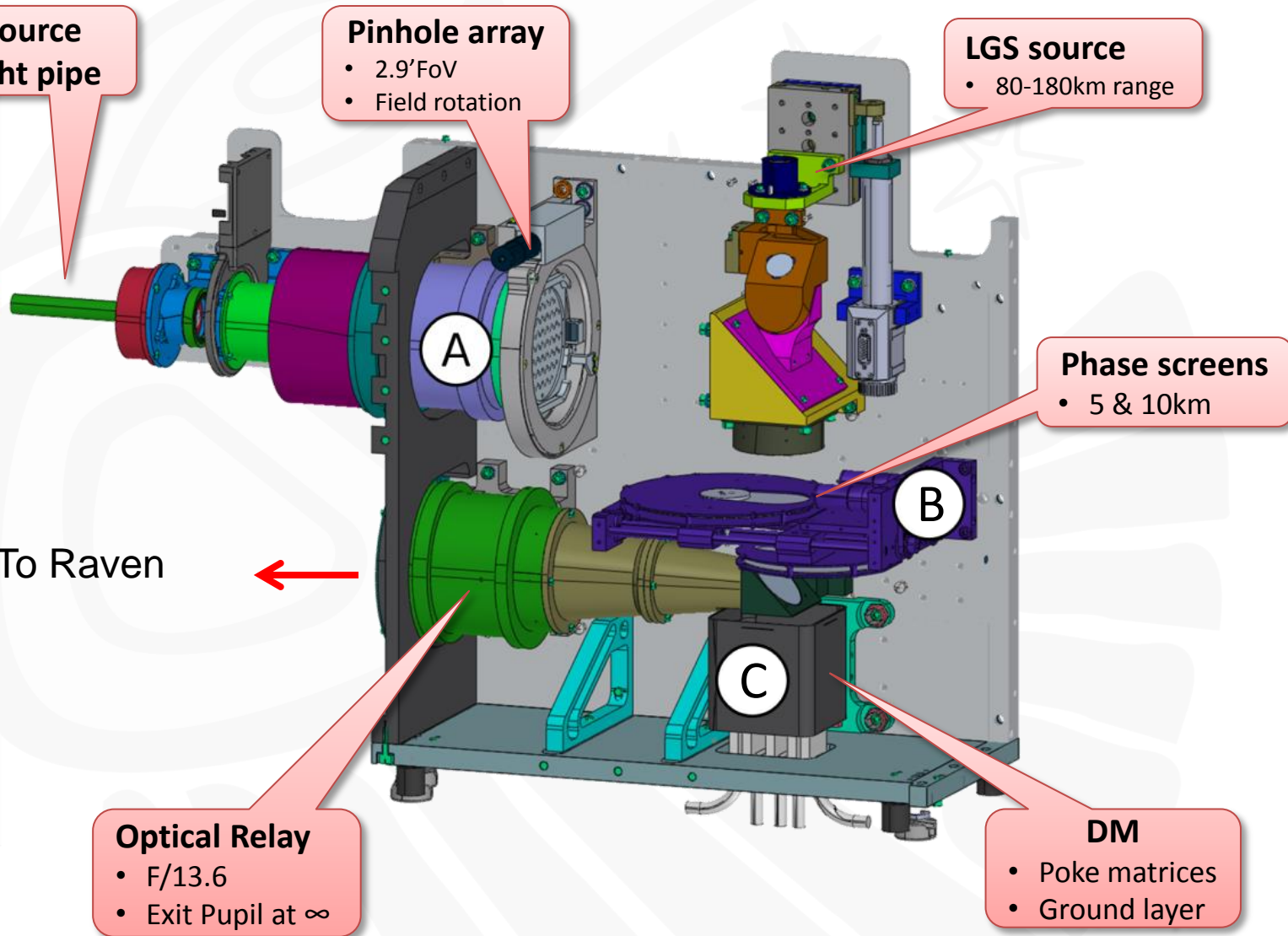
To Raven

Optical Relay

- F/13.6
- Exit Pupil at ∞

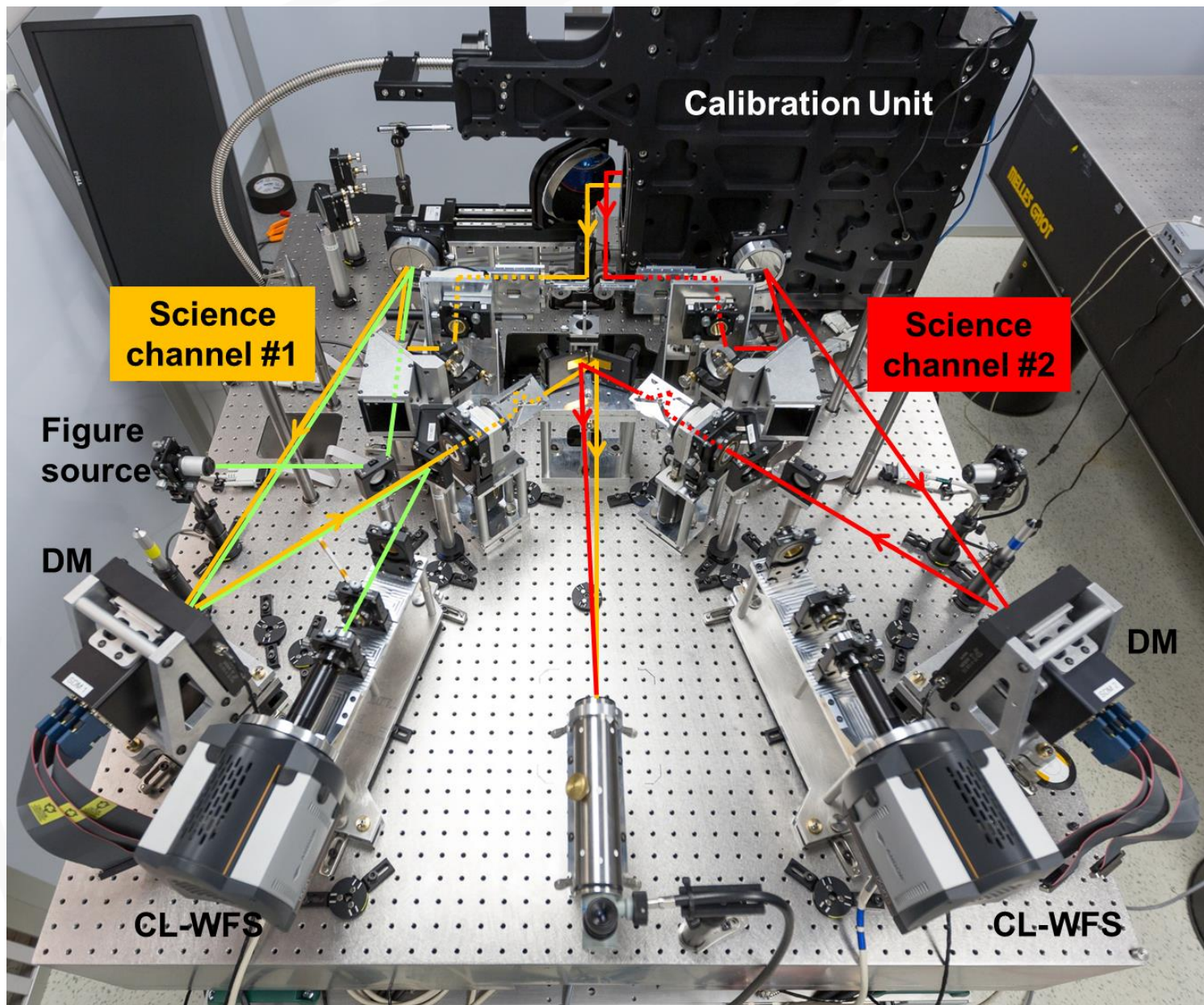
DM

- Poke matrices
- Ground layer





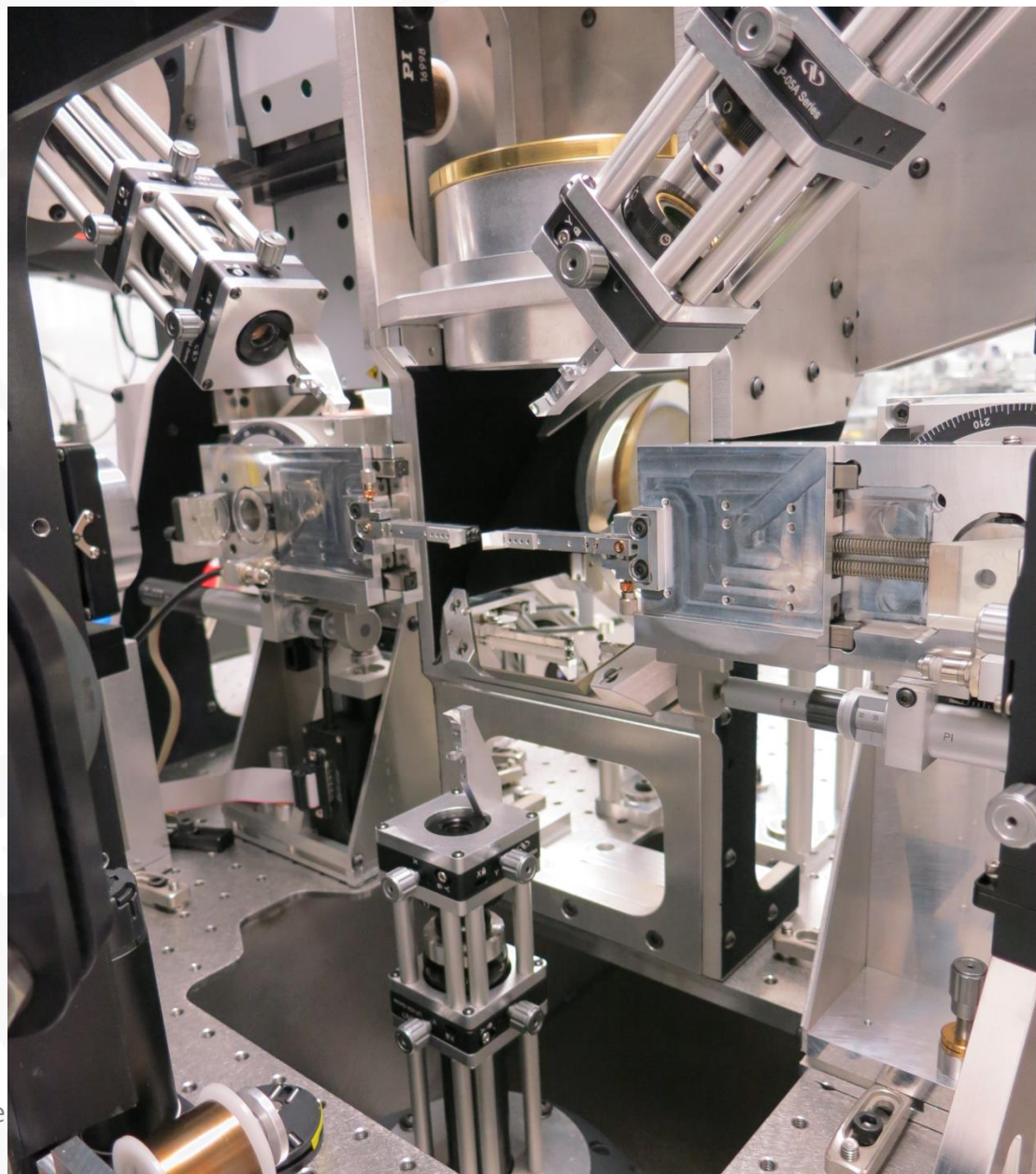
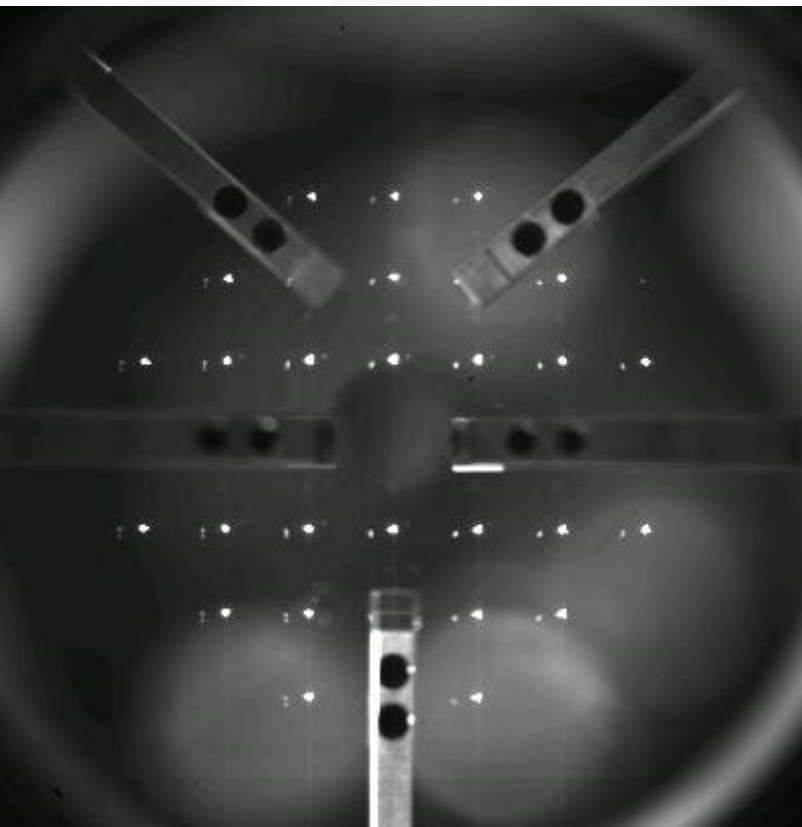
The 2 science channels implemented





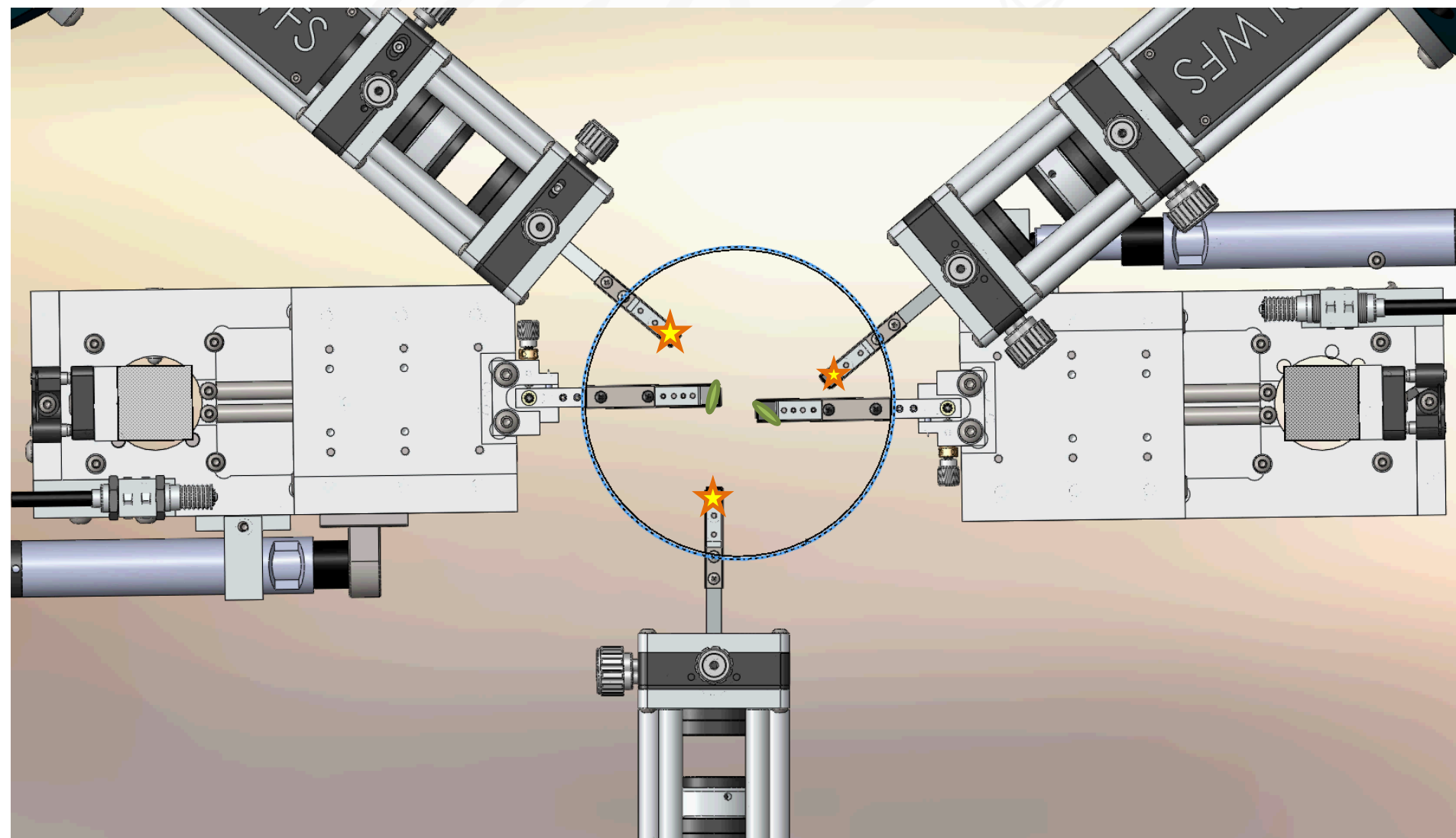
Raven pick-off mirrors

Acquisition camera image





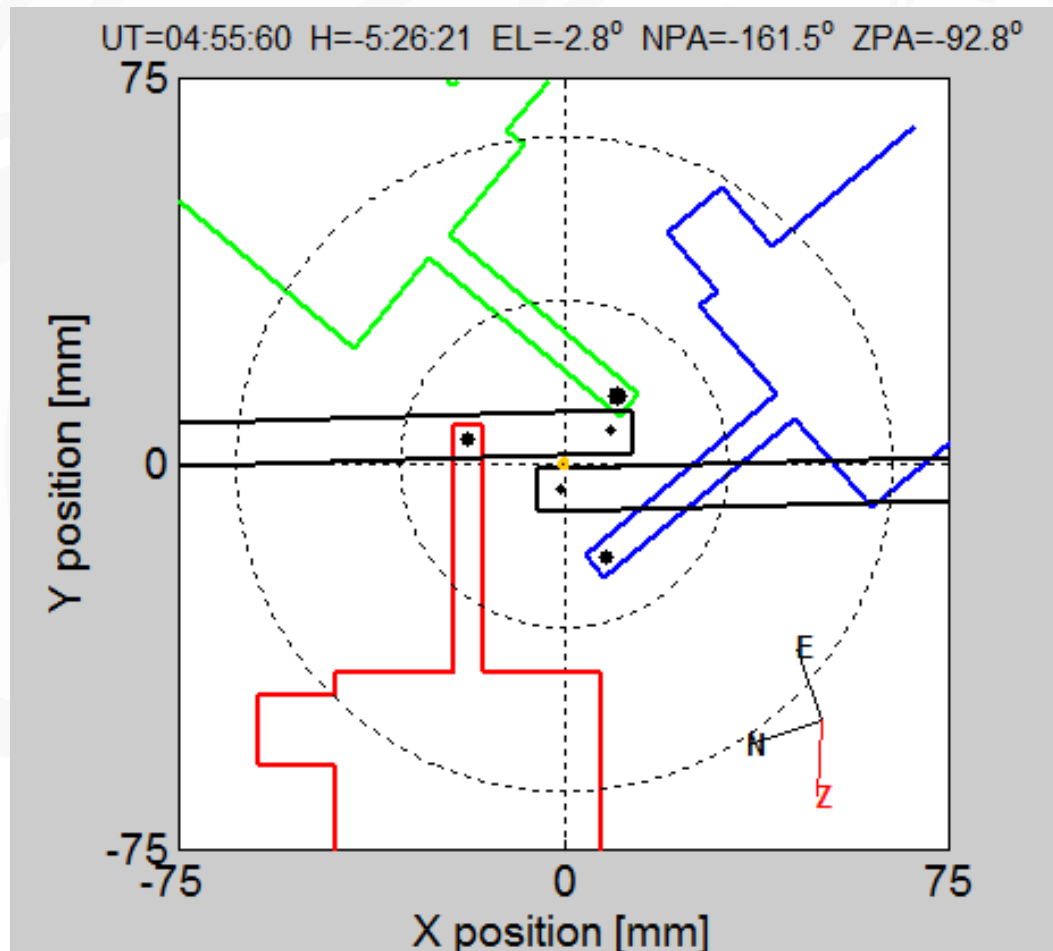
Field rotation tracking





Simulation for a real science case

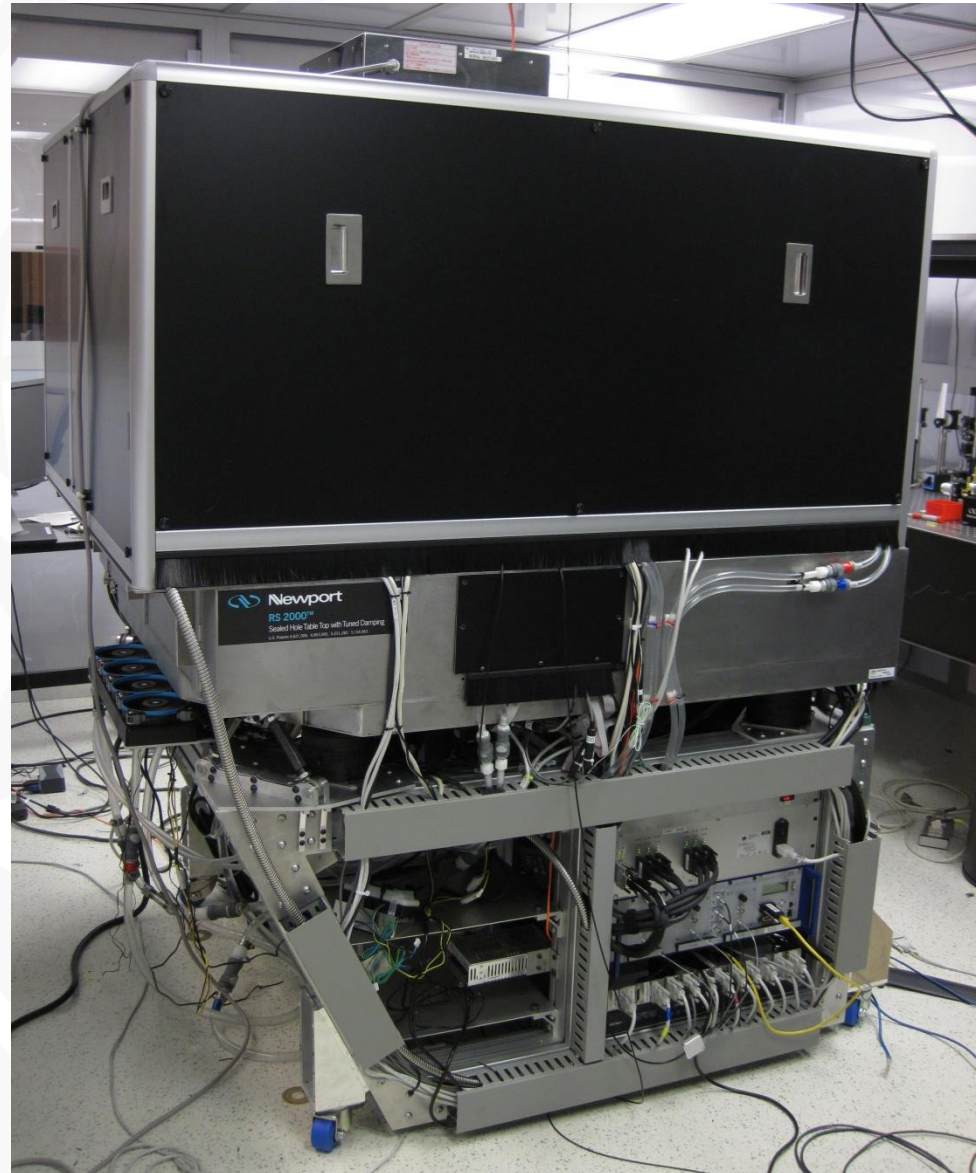
Example for Stanek field on 20/06/2013





Hardware completed at UVic (July 2012)

- All the opto-mechanical components are installed and aligned
- Most of the electronics are mounted in the frame under the bench

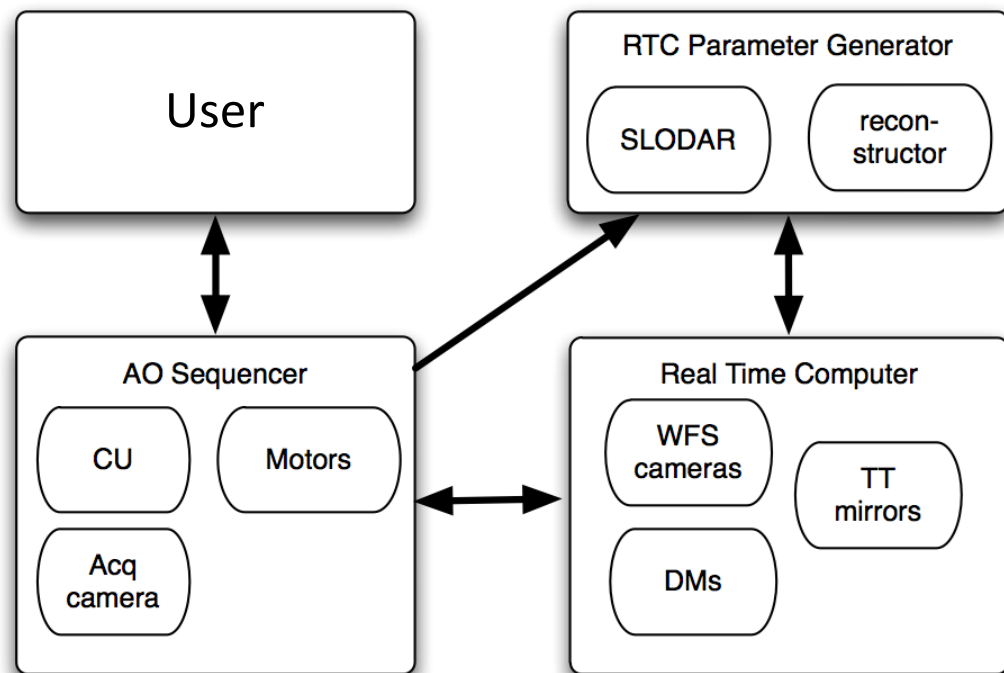




Software architecture

The Raven Software is divided into 3 subsystems:

- The **AO Sequencer (AOS)** sets up the system and controls all non-real-time hardware, provides a user interface.
- The **Real-Time Computer (RTC)** processes the WFS pixels and generates the DM commands.
- The **RTC Parameter Generator (RPG)** updates the tomographic reconstructor as the conditions change using the WFS data (SLODAR or Learn&Apply).





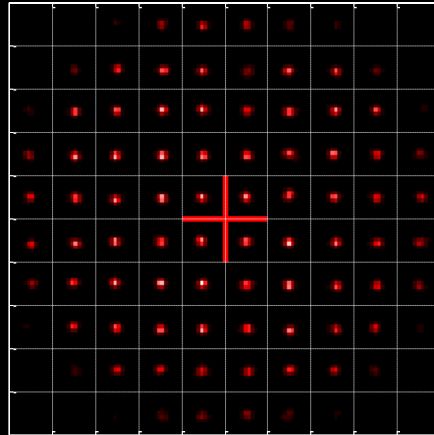
RTC supports different AO modes

AO mode	Description
SCAO	Single Conjugated AO: Close loop on bright science target
MOAO	Multi-Object AO: NGS (+LGS) WFSs feed tomographic reconstructor, then feed the DM (open loop)
GLAO	Ground-layer AO: WFS slopes are averaged and sent to the DM (open loop)
HP MOAO + LP SCAO	High-Pass-filtered MOAO + Low-Pass-filtered SCAO: <ul style="list-style-type: none">• Fast turbulence is corrected in open-loop MOAO,• Slow turbulence & quasi-static aberrations are corrected in close-loop w/ the CLWFS running at low frame rate on a possibly faint compact science target.

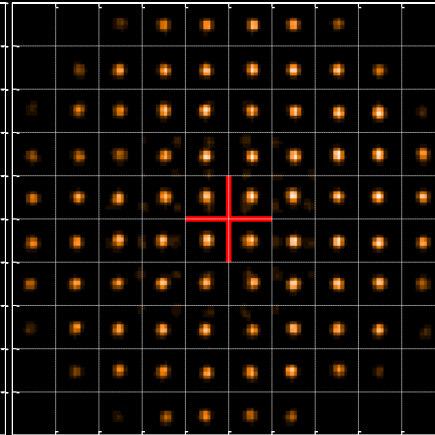


AOS user interface in Matlab

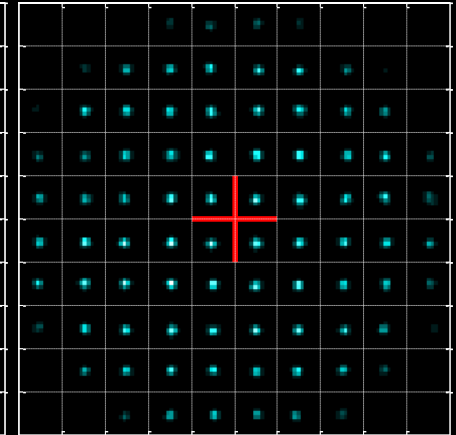
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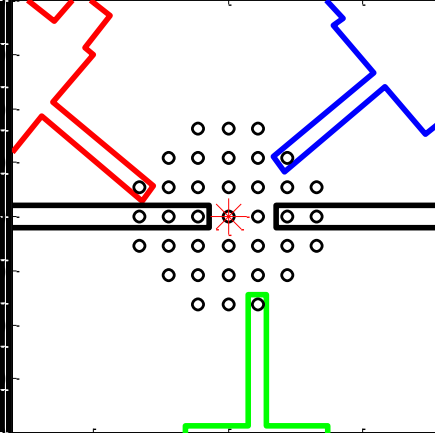
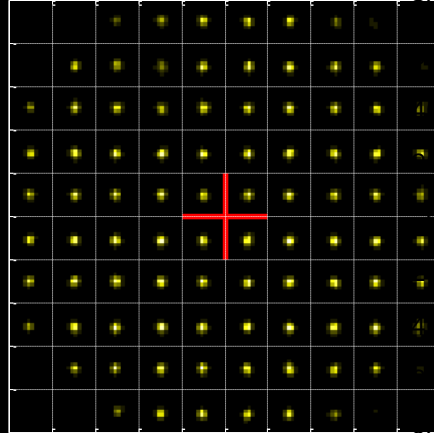
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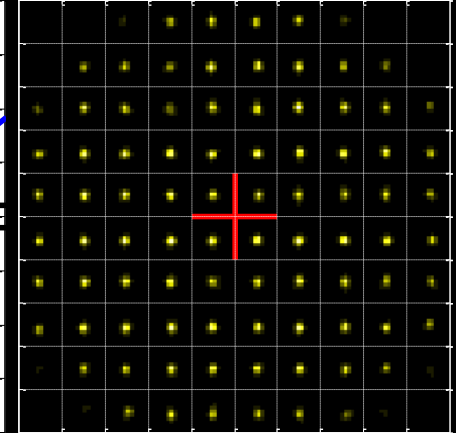
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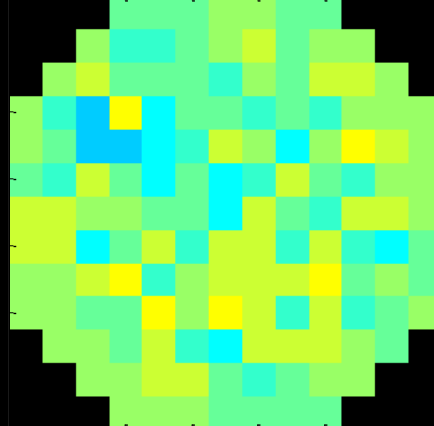
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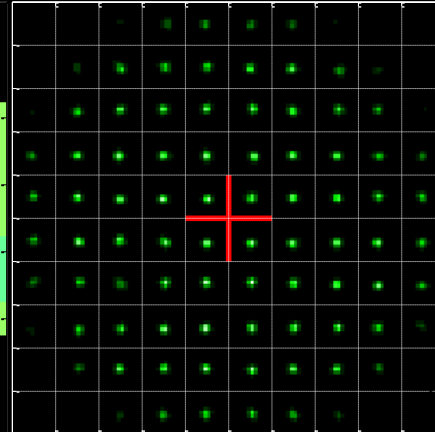
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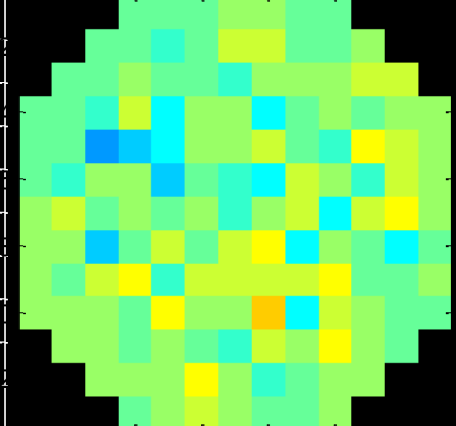
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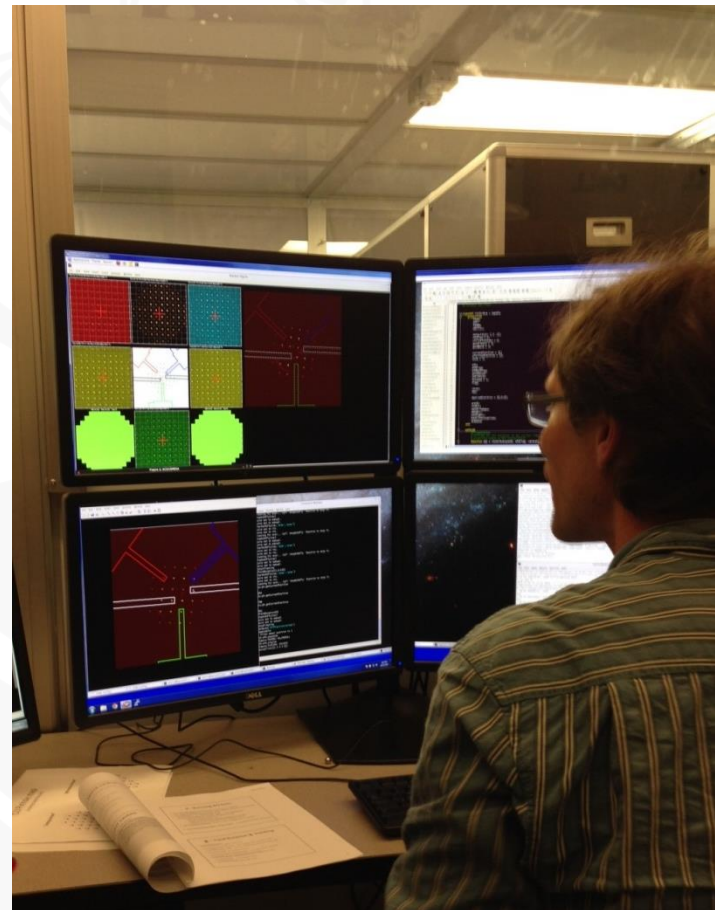


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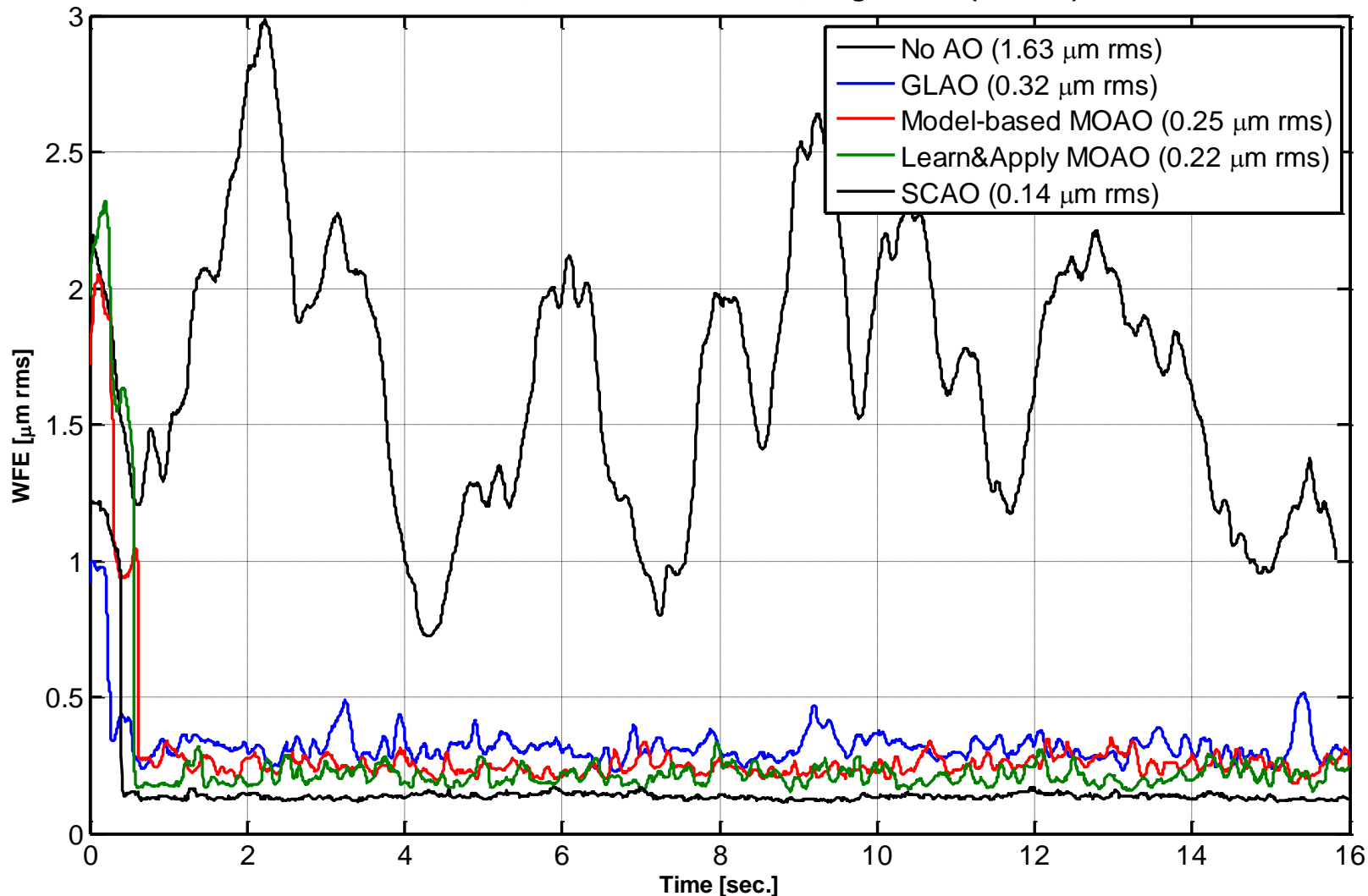
Raven Acceptance Review at UVic, 26 Nov. 2014





Lab Results obtained at UVic

CL-WFS data, Wide asterism + LGS, Bright stars ($R=9\sim 10$)

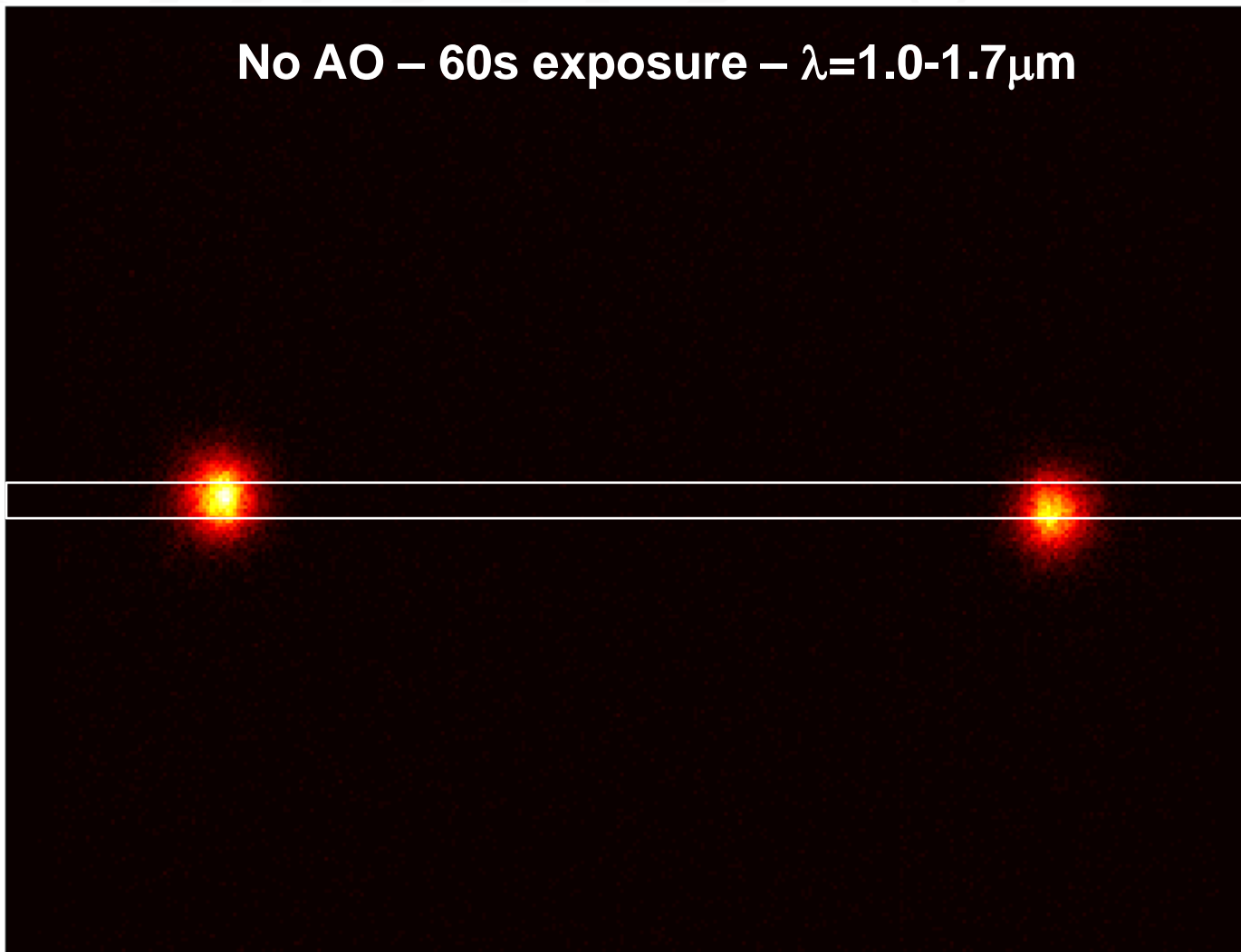




Science camera long-exposure images

No AO – 60s exposure – $\lambda=1.0\text{-}1.7\mu\text{m}$

0.14" \updownarrow





Science camera long-exposure images

MOAO – 60s exposure – $\lambda=1.0-1.7\mu\text{m}$

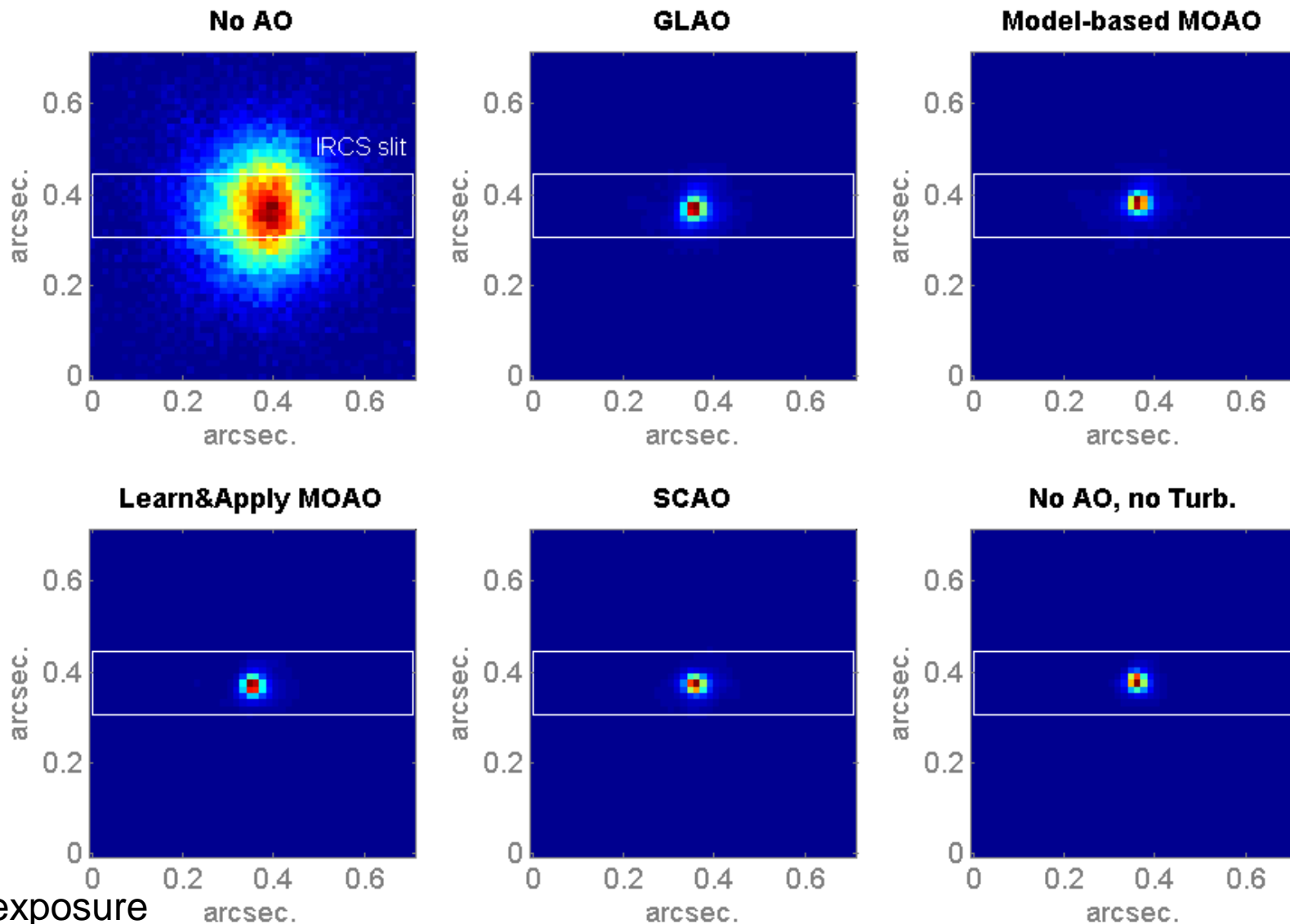
0.14" \updownarrow





Science Camera long-exposure images

Bright & Wide Asterism + LGS





Conclusions

Current performance meet science requirements

- *In agreement with error budget*
- *Good image correction: Strehl > 20-50%*
- *Limiting magnitude $R \geq 14.25$ w/ CoG*

Upgrades & expected improvements

- *Correlation centroiding (+1 mag.)*
- *Predictor (+2mag)*



Conclusions



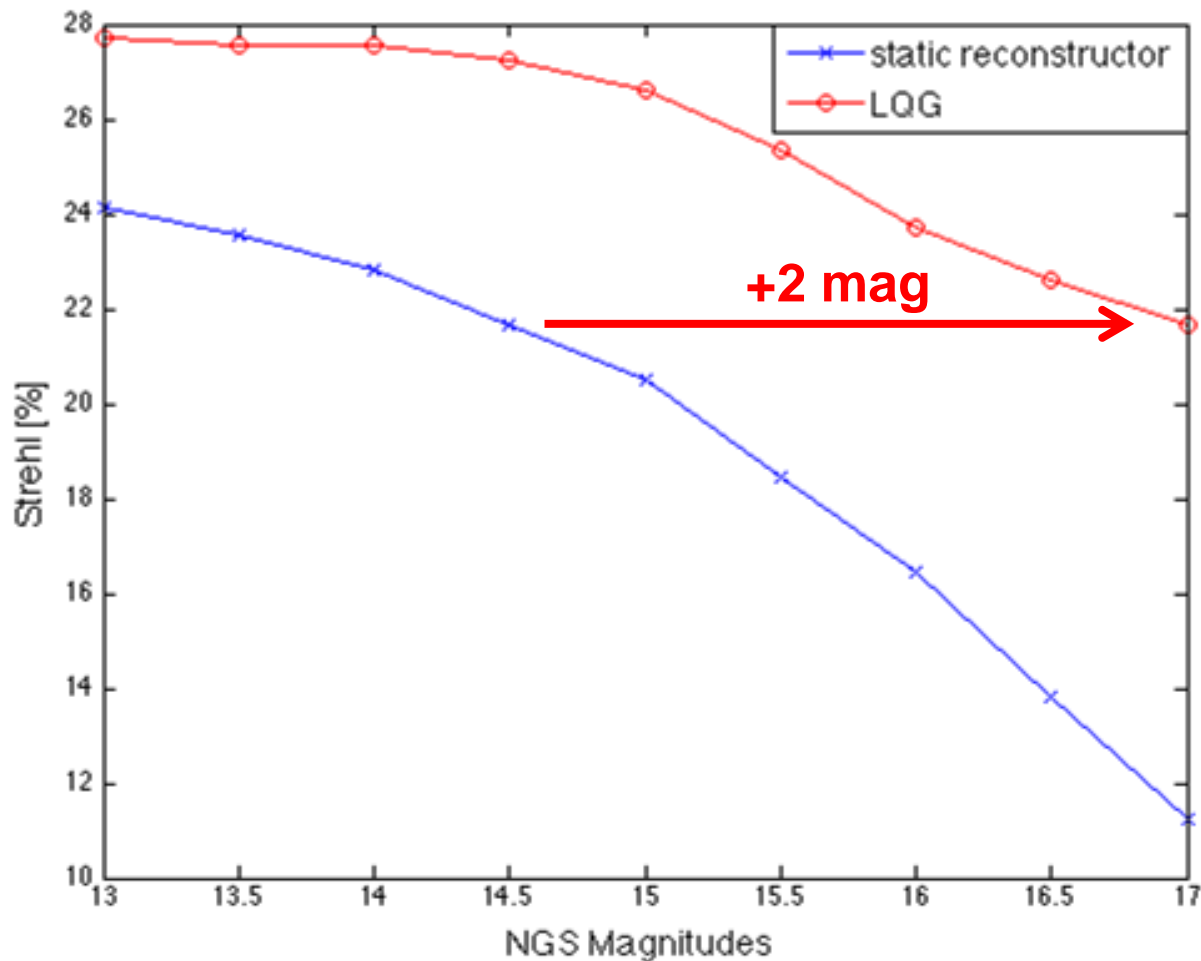
Current performance

- In agreement with e
- Good image correcti
- Limiting magnitude



Upgrades & expe

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Conclusions

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Upgrades & expected improvements

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Science instrument ready to go on sky:

- *Engineering nights in May and Aug 2014*
- *Hoping for science nights in S14B and S15A*
- *Little risks for science cases as GLAO meets science requirements too.*



Raven arrival (Jan 6th)

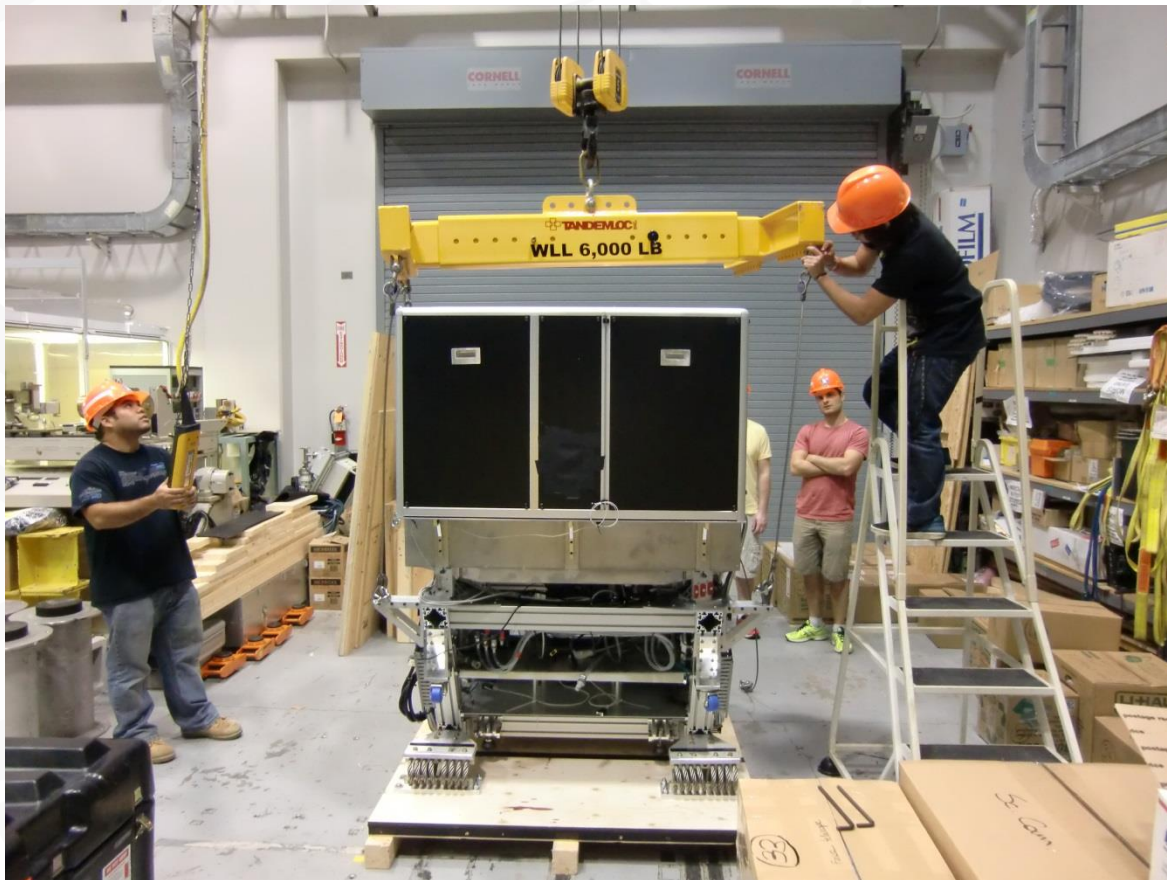
Matson SS Maui from Seattle to Honolulu



Photos: Shin Oya

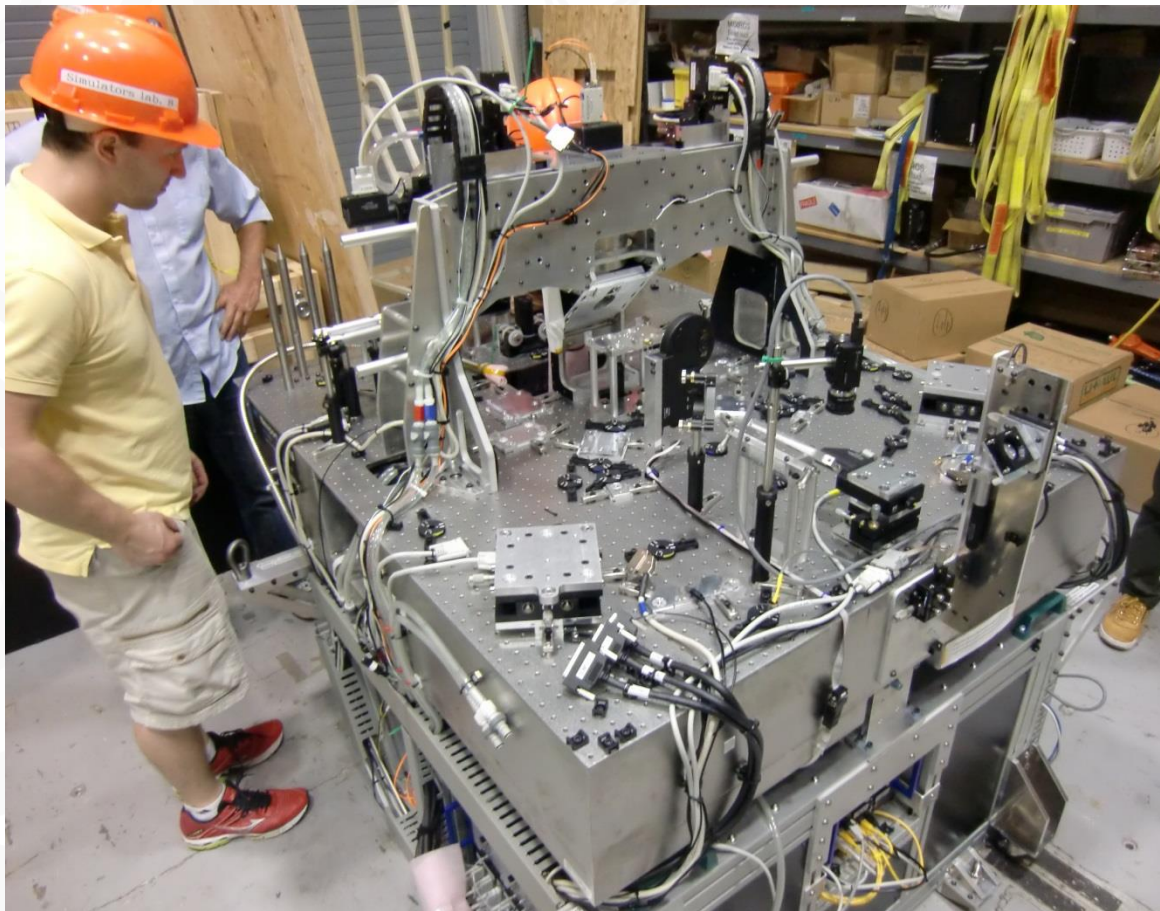


Raven in the SimLab (Jan 7th)





Raven in Simlab (Jan 7/8)





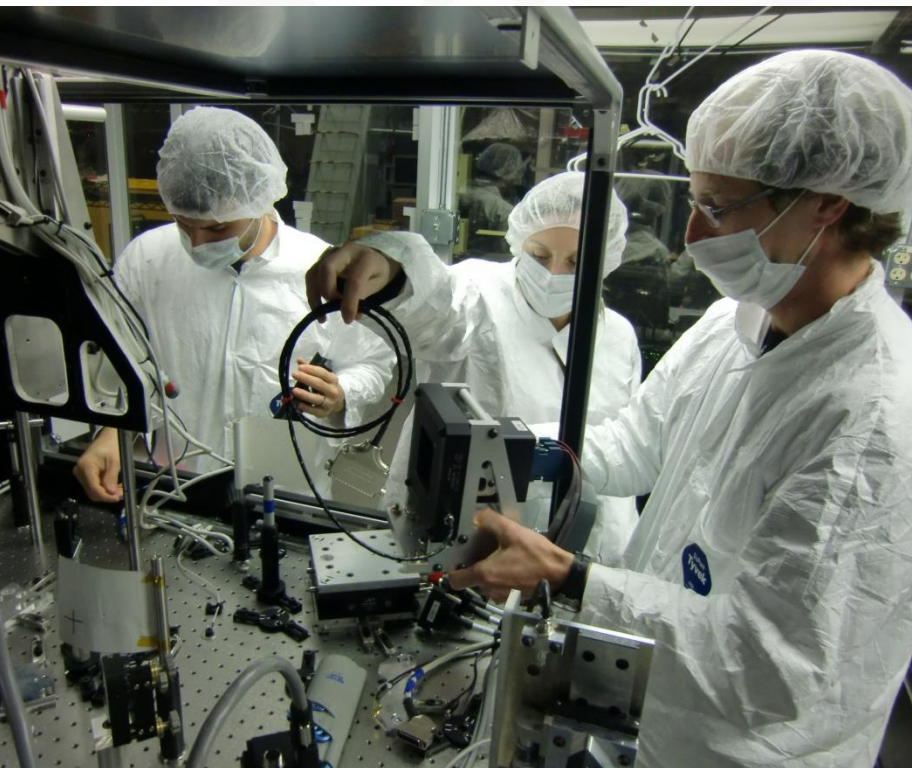
Raven in Simlab (Jan 7/8)



Photos: Shin Oya



Alignment in SimLab clean room





2014 Schedule

- ✓ Jan : Raven integration and alignment in the SimLab
- Feb-Mar :
 - Software upgrades & consolidation
 - Final tests
 - Prepare observation plan
- Mid-Apr : Raven ships to summit in BSIT truck
- **May 13/14 : First engineering nights**
- Aug : 2nd run of engineering nights



Thank you

