

# **Development of an integral field unit for a NIR-MOS camera: SWIMS**

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@ Hokkaido university**

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# Collaborator



## ✓ **SWIMS team in TAO project**

- Kentaro Motohara (PI)
- Masahiro Konishi
- Hidenori Takahashi
- Natsuko Kato
- Ken Tateuchi
- Soya Nishijima

## ✓ **NAOJ-ATC\***

- Shinobu Ozaki

## ✓ **Kyoto-Sangyo Univ.**

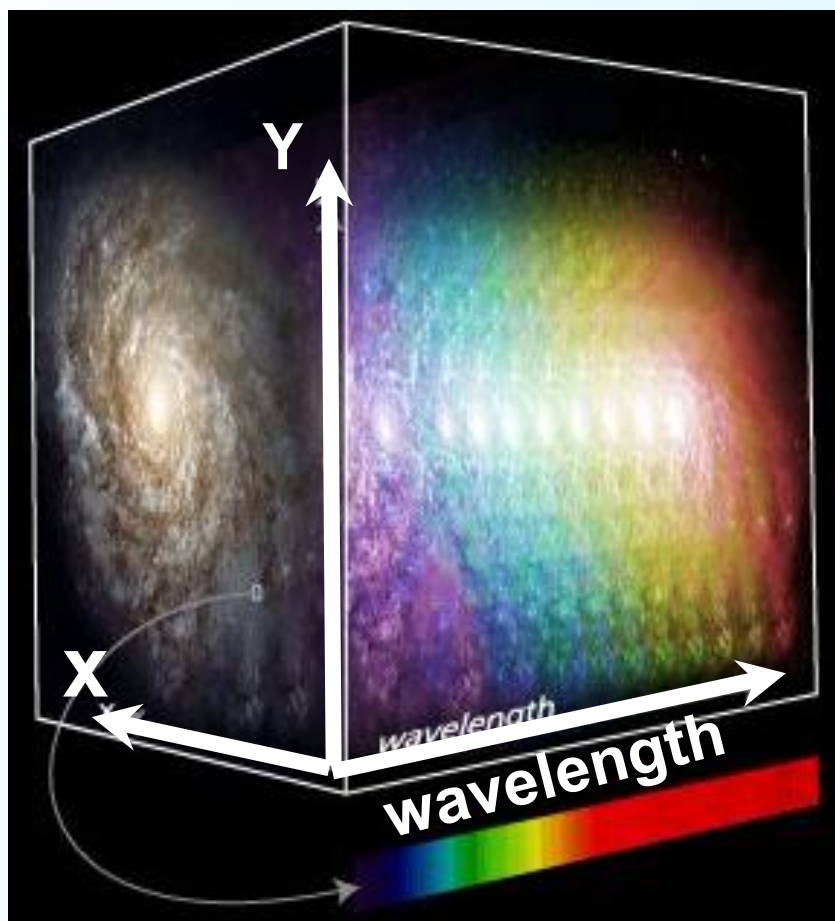
- Tomohiro yoshikawa

\*Advanced Technology Center

- ✓ **Introduction**
  - What is an IFS?
  
- ✓ **Development of SWIMS-IFU**
  - concept
  - mechanical/optical design
  
- ✓ **Towards the future**
  - next generation instruments with IFU
  
- ✓ **Summary**

## What is an integral field spectrograph?

✓ IFS enables to obtain spectra over a 2D field- of- view



### “Data Cube”

2D spacial info. = (RA, DEC)

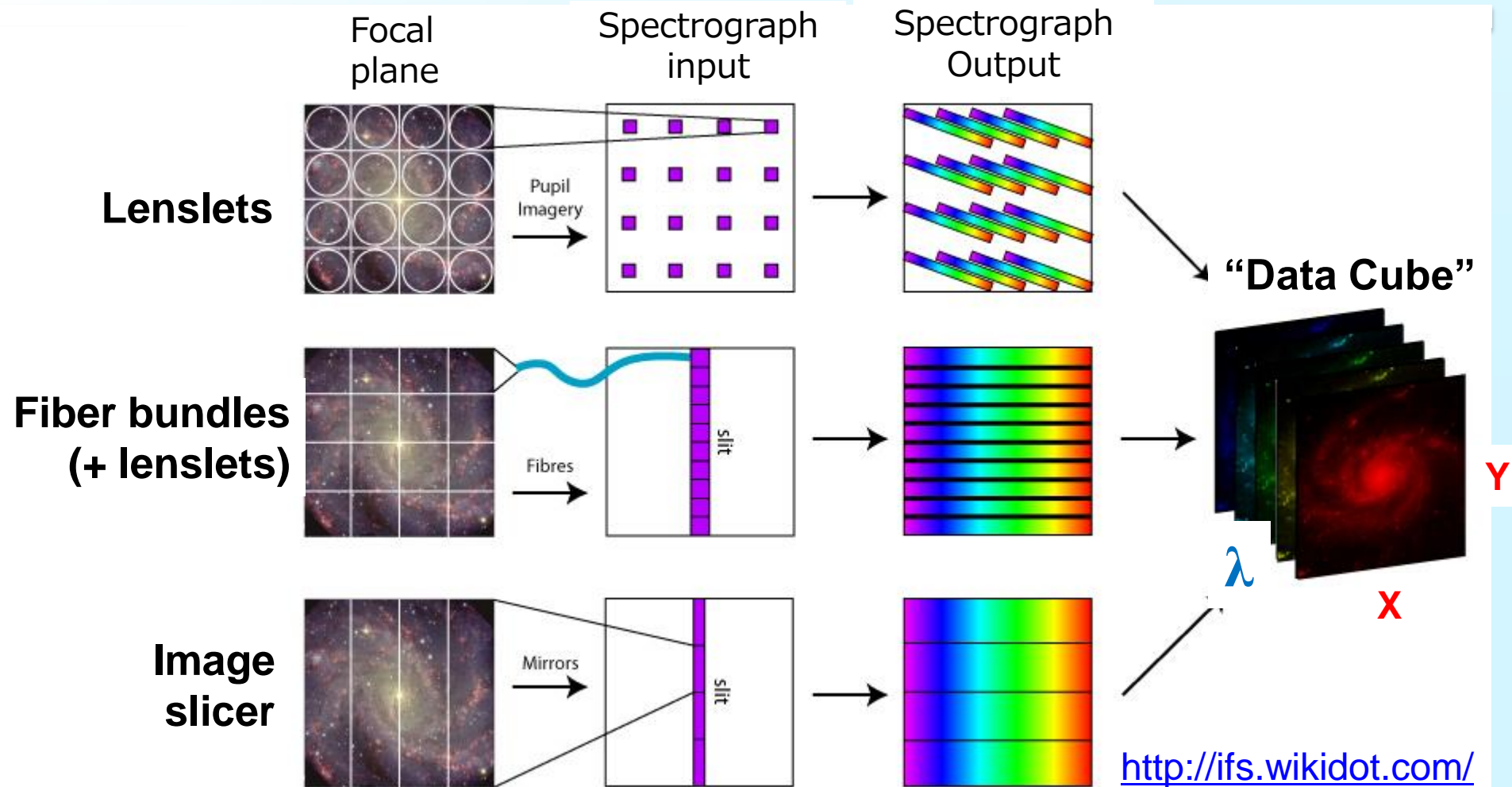
+

1D spectral info. =  $\lambda$  or  $\nu$

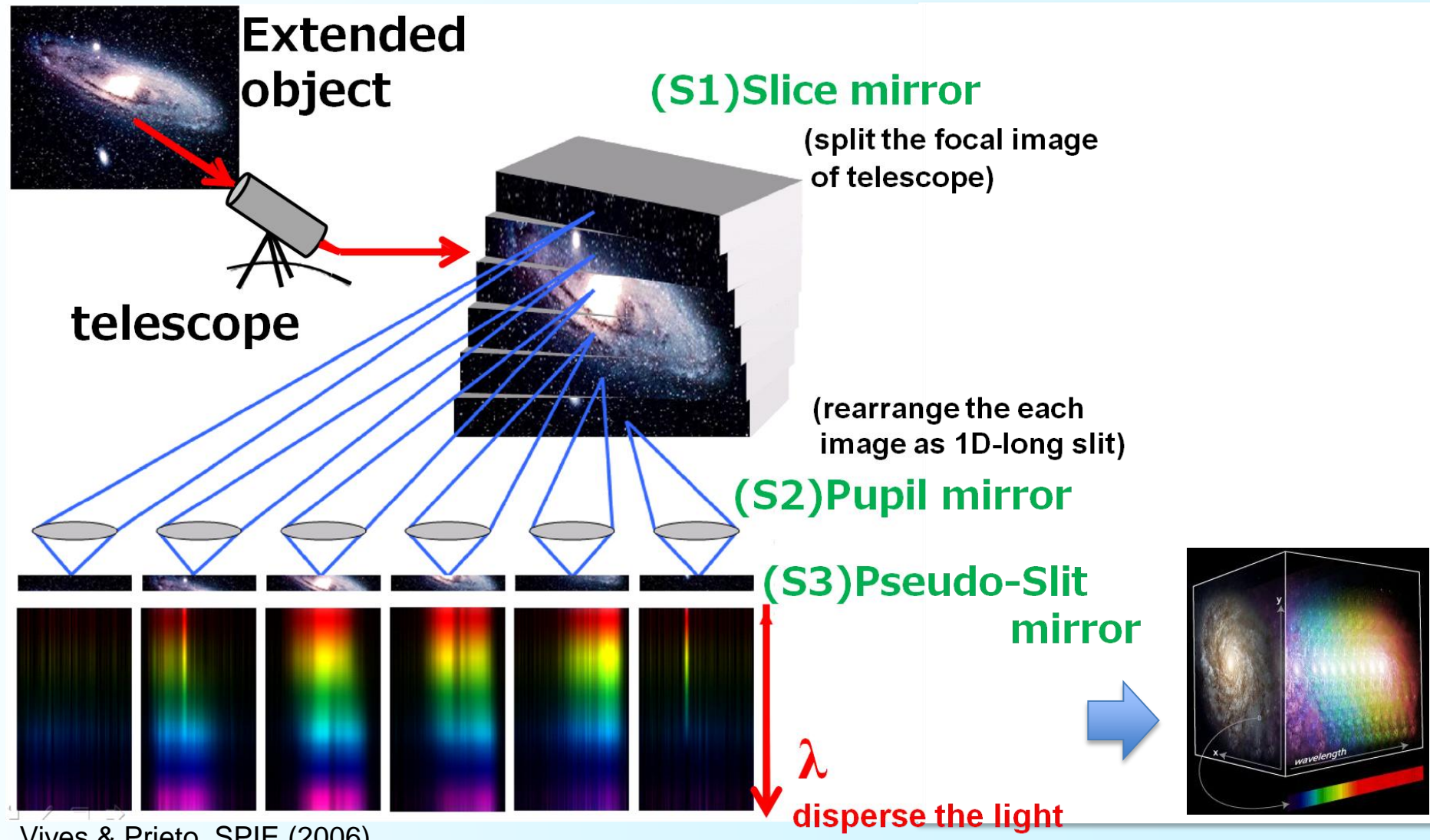
<http://sciencewise.anu.edu.au/>

# Main IFS techniques

## The 3 types of Integral Field Unit (= IFU)



# The concept of Advanced image slicer



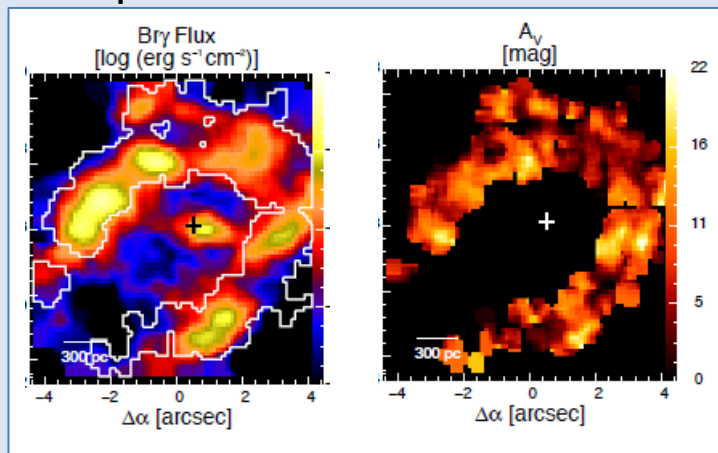
Vives & Prieto, SPIE (2006)

# near-IR Integral Field Spectroscopy

## ✓ The need for near-IR spectroscopy

- (1) Less affected by dust obscuration (e.g. dusty starbursts)
- (2) At  $z > 1$  most of the key spectral features are redshifted in the near-IR

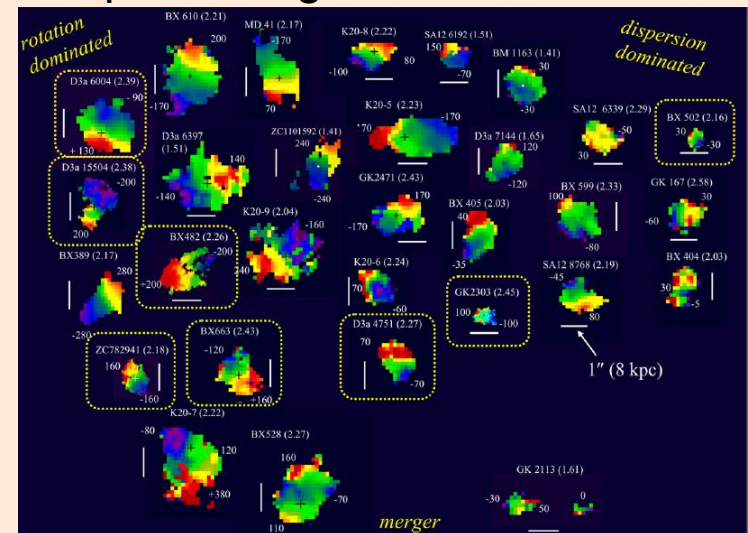
### • Samples in the local universe



LIRG ( $z=0.01$ ) with high spatial resolution ( $\sim 0.3$  kpc)

J. Piqueras Lopez et al. (2013)

### • Samples at high redshift



Forster Schreiber et al. (2009)

# near-IR Integral Field Unit

✓ current near-IR IFUs of the 8-10m class telescopes

telescope	VLT (8.1m)	Keck (10m)	Gemini (8.1m)
NIR-IFU	<b>SINFONI</b>	<b>OSIRIS</b>	<b>NIFS</b>
IFU-type	<b>slicer</b>	<b>lenslets</b>	<b>slicer</b>
$\lambda$ [ $\mu\text{m}$ ]	<b>1.1-2.45</b>	<b>1.0-2.4</b>	<b>0.94-2.4</b>
$R=\lambda/\Delta\lambda$	<b>2000-4000</b>	<b>~3800</b>	<b>5000-6000</b>
Spatial element size	<b>0.0125" x 0.025" 0.050" x 0.10" 0.125" x 0.250"</b>	<b>0.020" , 0.035" 0.050" , 0.100"</b>	<b>0.104"x0.04"</b>
Field of view	<b>0.8" x 0.8" 3" x 3" 8" x 8"</b>	<b>0.32" x 1.28" - 3.2" x 6.4"</b>	<b>3" x 3"</b>



# Conceptual design of SWIMS-IFU

## ✓ SWIMS-IFU

- advanced image slicer IFU
- compact & lightweight  
→ enable to install into MOS camera as a “mask-slit”

parameters of  
SWIMS-IFU

$\lambda$ [um]	FoV	Slice width	No. of slice
0.9 – 2.5	14" × 10"	0.4"	26

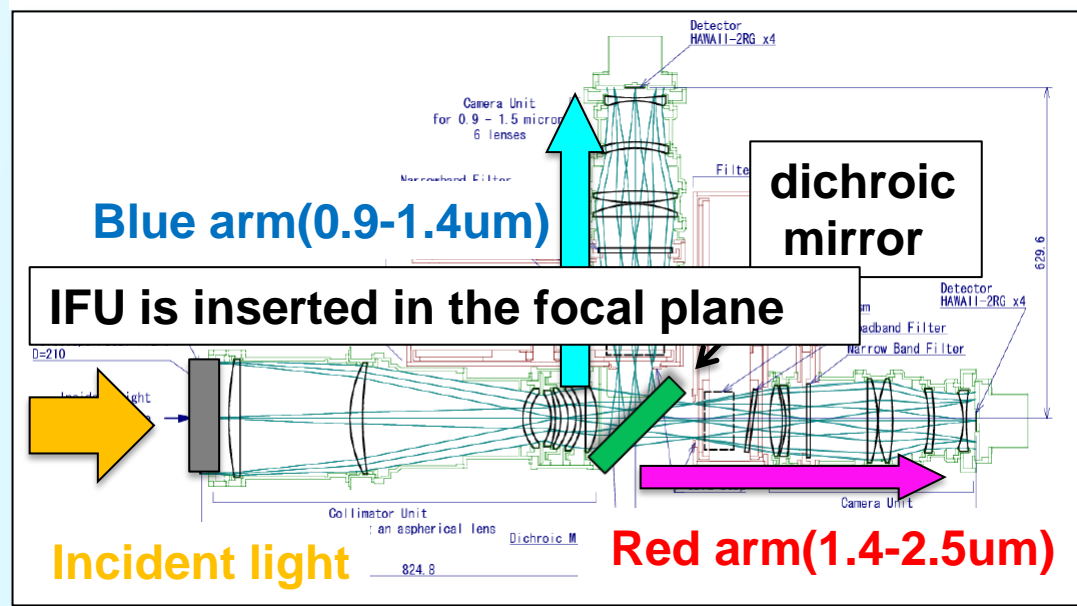
# NIR-MOS camera: SWIMS

✓ SWIMS = **S**imultaneous-color

↓ **W**ide-field  
**I**nfrared  
**M**ulti-object  
**S**pectrograph

SWIMS is capable of

- **wide-field** imaging(6.6' x 3.3') & spectroscopy(2.8' x 3.3')
- covering the entire NIR spectra (0.9- 2.5 um) in a single exposure



M. Konishi et al. Proc. SPIE, 8450-144 (2012)

“Development of an integral field unit for SWIMS”

# Development of SWIMS-IFU

✓ What are problems ??

(1) IFU is mounted inside the SWIMS slit slide mechanism

➡ **“compact & lightweight” design is required !**

*But such a optical solution can be really exist...?*

(2) IFU is placed under a cryogenic environment

➡ **Metal processing mirrors are desirable !**

*How can we achieve required surface accuracy...?*

# Development of SWIMS-IFU

✓ What are problems ??

(1) IFU is mounted inside the SWIMS slit slide mechanism



**“compact & lightweight” design is required !**

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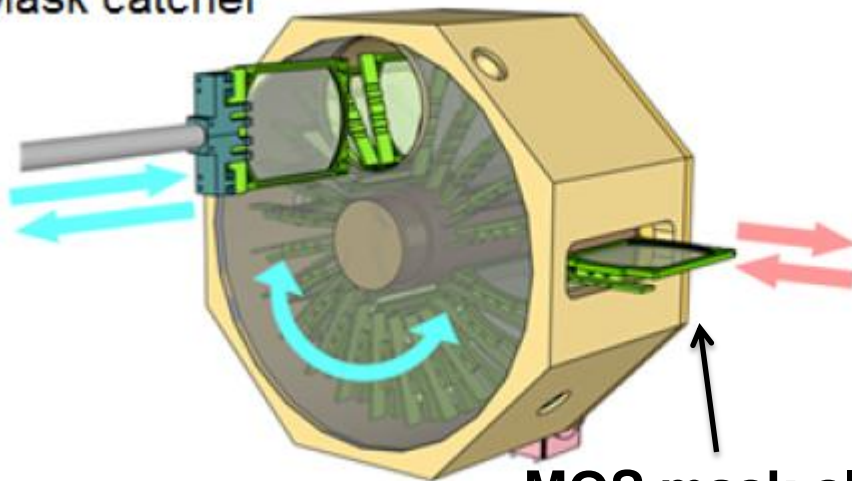
**Metal processing mirrors are desirable !**

*How can we achieve required surface accuracy...?*

# Mechanical design

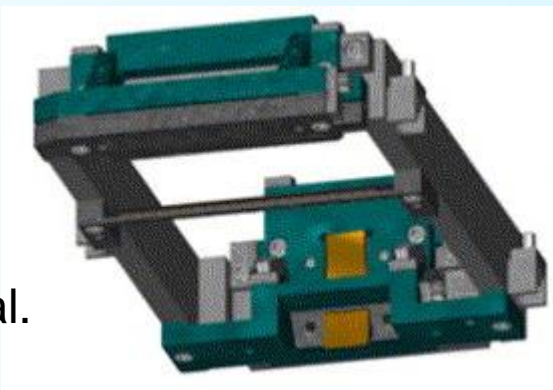
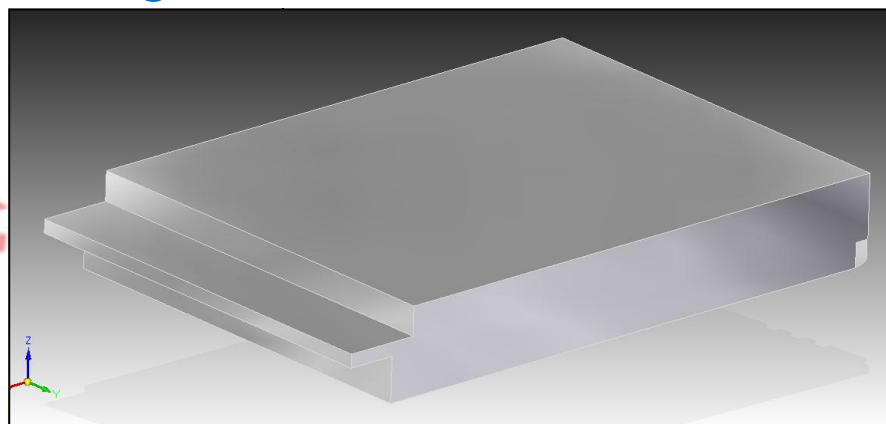
✓ very compact & lightweight unit

Mask catcher



MOS mask-slit

## Integral Field Unit



Size	$170 \times 140 \times 40 \text{ mm}^3$
Mass	$< 700\text{g}$

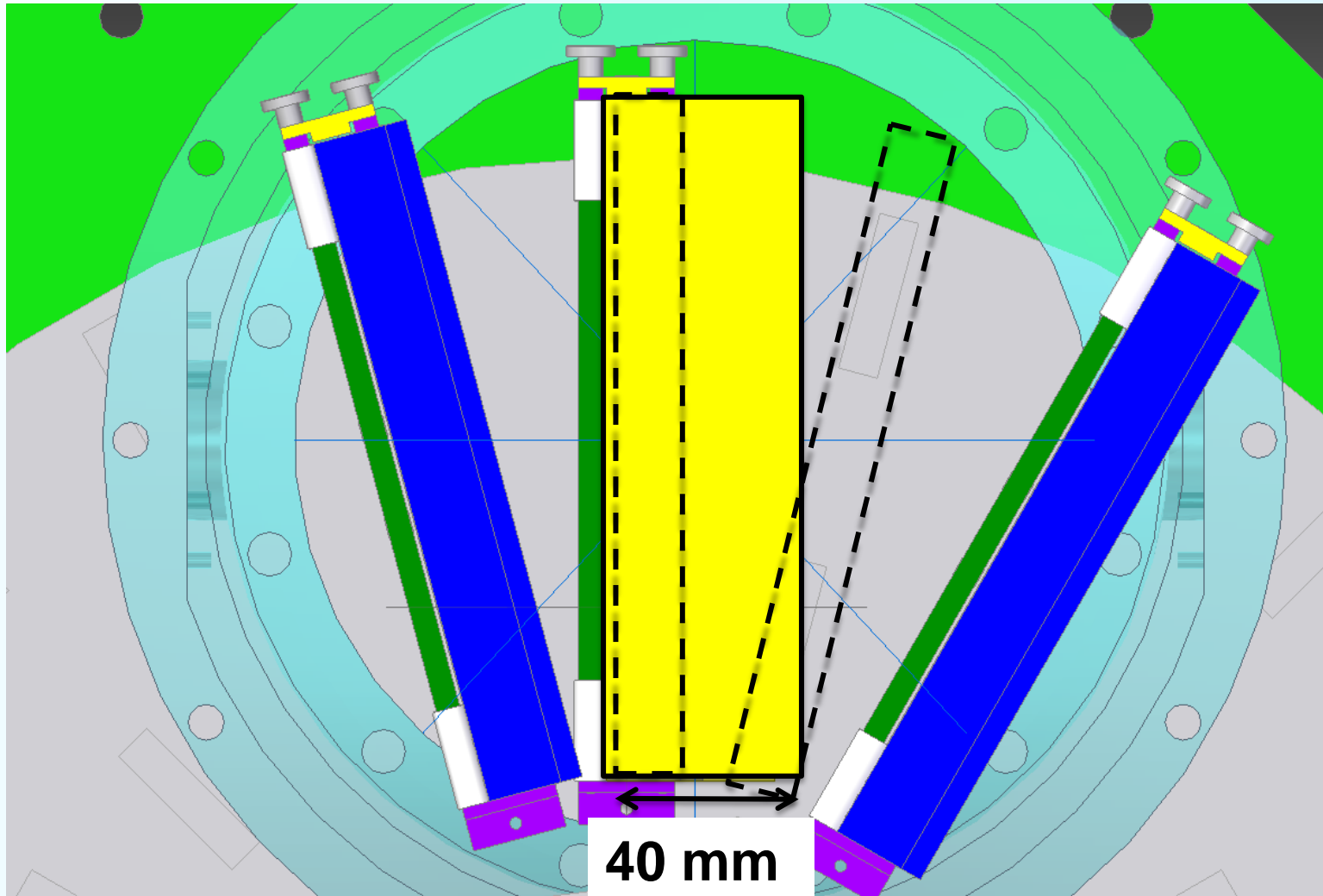
<GNIRS/GMOS>

Size:  $200 \times 100 \times 100 \text{ mm}^3$

Mass:  $< 1000\text{g}$

Cornelis et al.  
(2006)

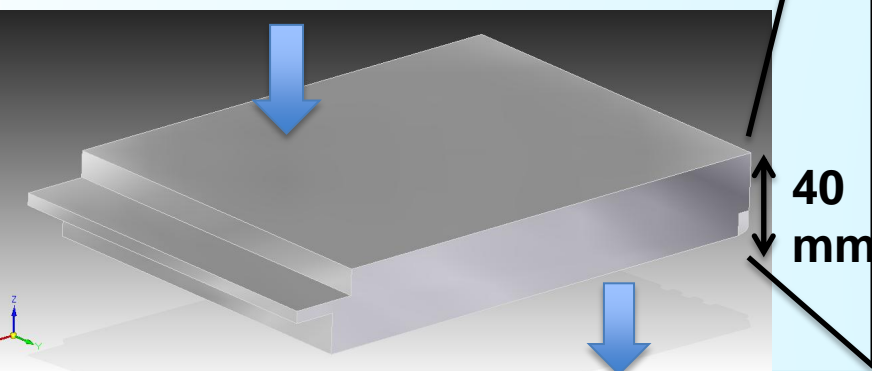
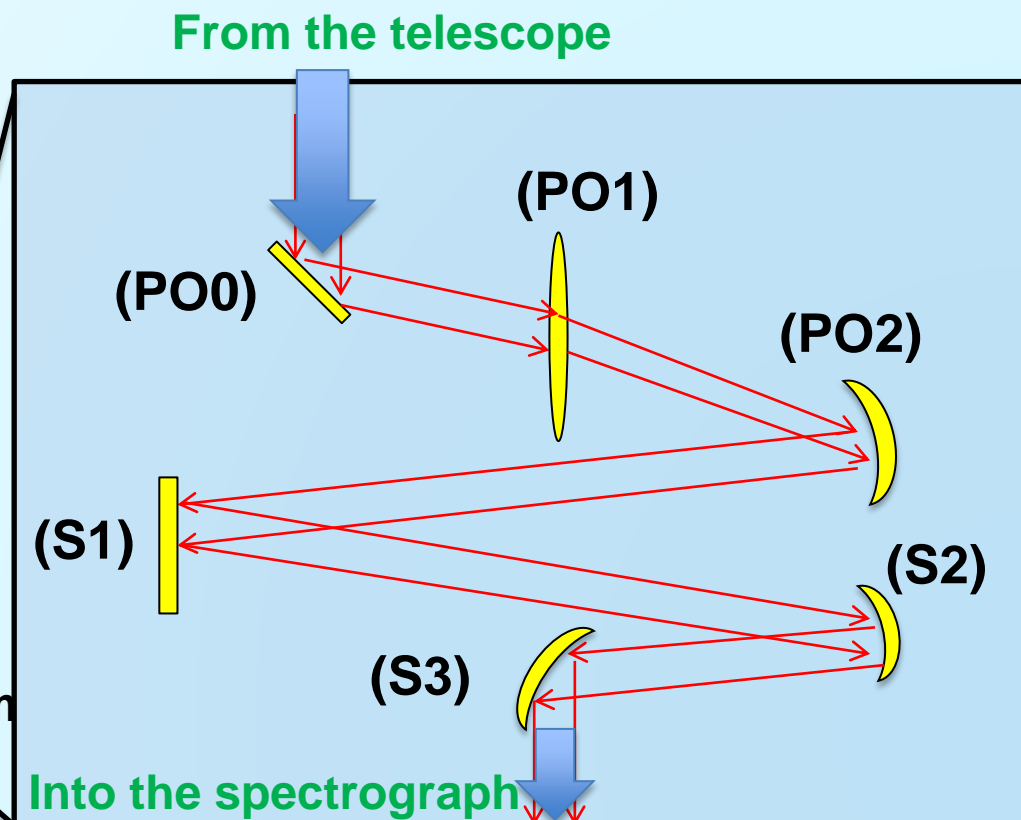
# Size limit to install into MOS camera



# Optical design

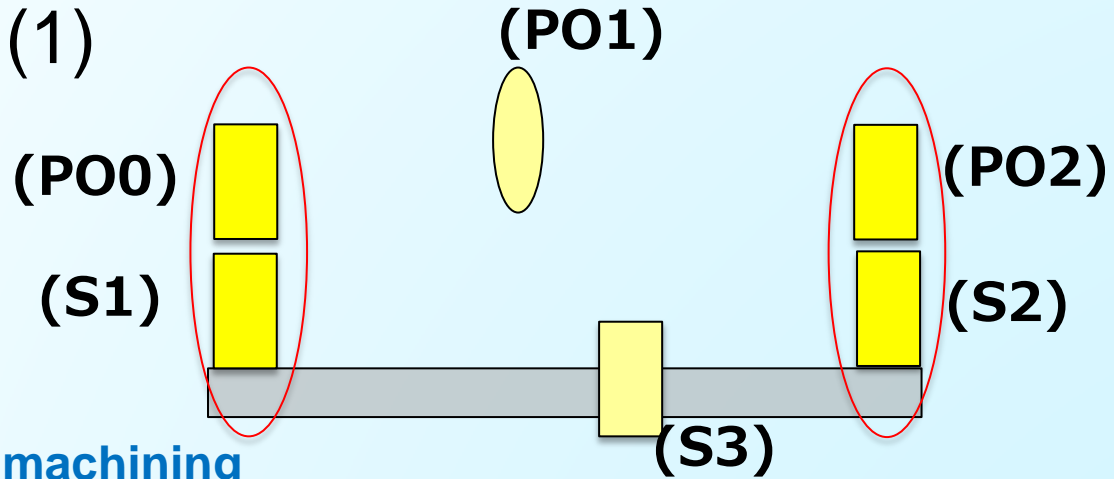
## ✓ Optical components of SWIMS-IFU

- (PO0) pick-off mirror
- (PO1) triplet lens
- (PO2) concave mirror
- (S1) Slice mirror
- (S2) Pupil mirror
- (S3) Pseudo-slit mirror



# Optical design

✓ Condition (1)

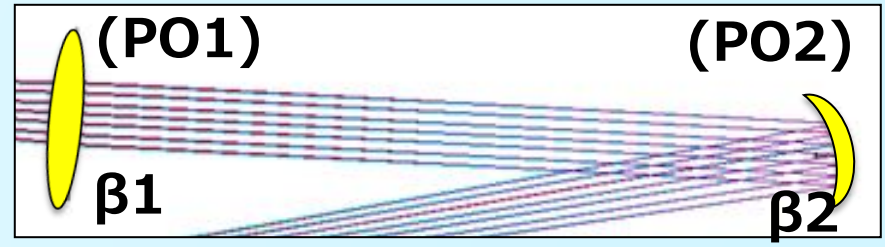


integrally machining  
 → can be achieved alignment accuracy in processing precision

✓ Condition (2)

collimated light between  
 (PO1) and (PO2)

PO1 & PO2 work as magnifier

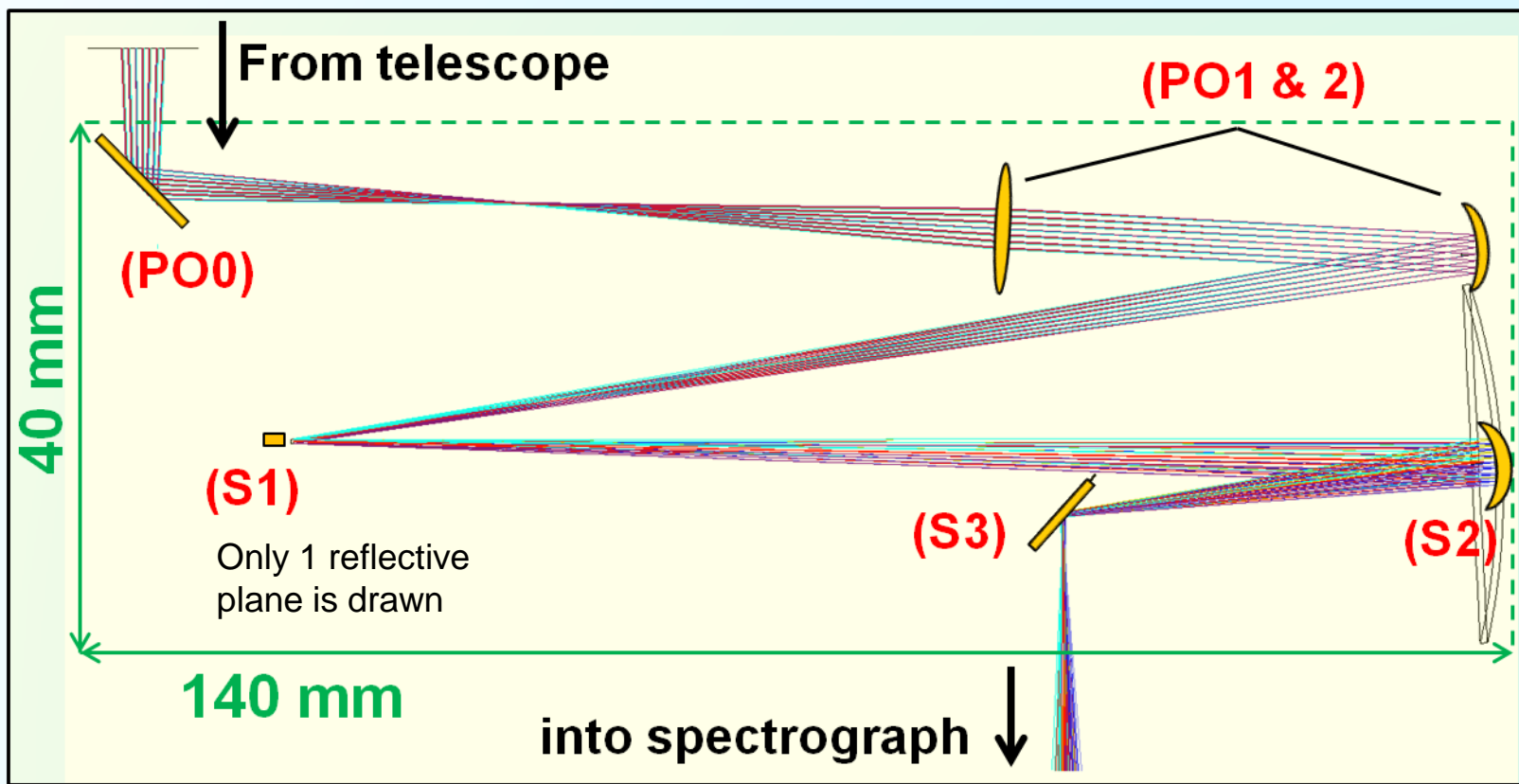


✓ Condition (3)  $\beta(\text{magnification}) = \beta_1 \times \beta_2 = 2.5$



# Optical design

- ✓ constructing the solution which satisfies the size limit



But in this case, each reflective surface of slice mirror has curvature...

# Development of SWIMS-IFU

## ✓ What are problems ??

(1) IFU is mounted inside the MOS slit-mask exchanger

➔ “compact & lightweight” design is required !

*But such a optical solution can be really exist...?*

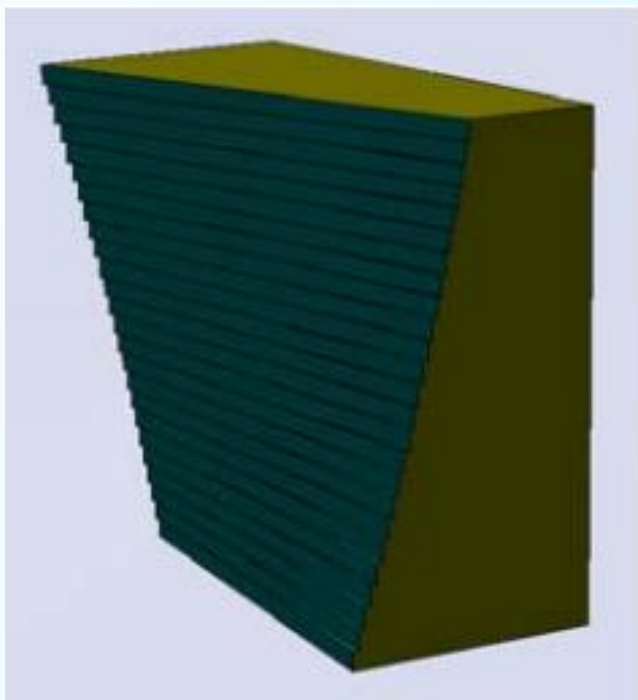
(2) IFU is placed under a cryogenic environment

➔ **Metal processing mirrors are desirable !**

*How can we achieve required surface accuracy...?*

# Fabrication of a slice mirror

Slice mirror (flat)



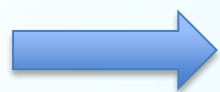
Processing method using glass can achieve high accuracy.

But it is

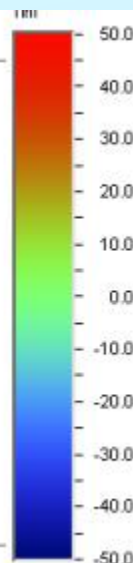
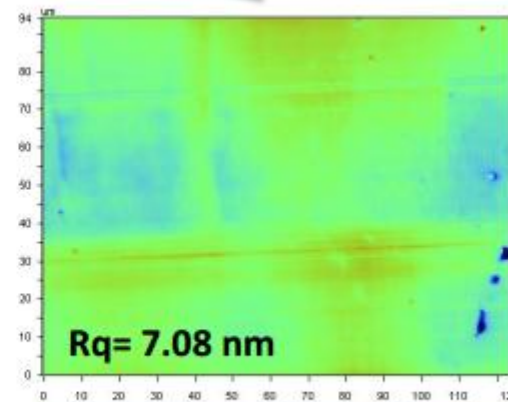
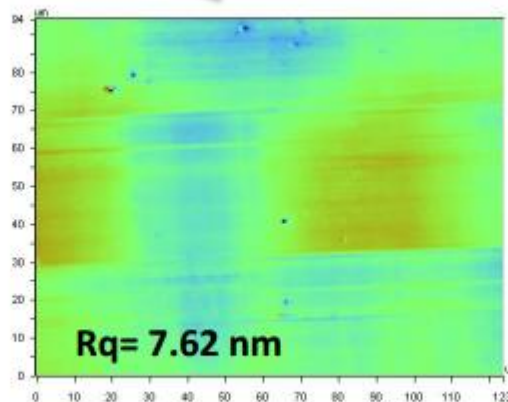
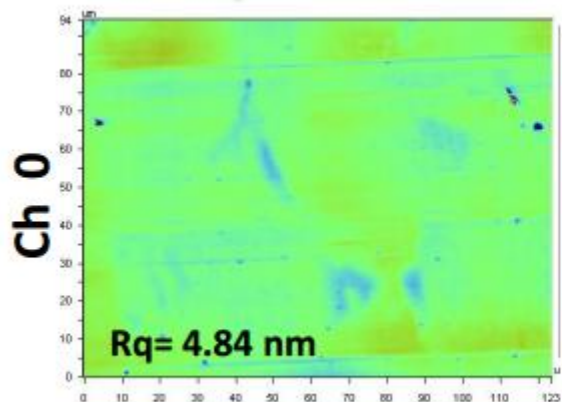
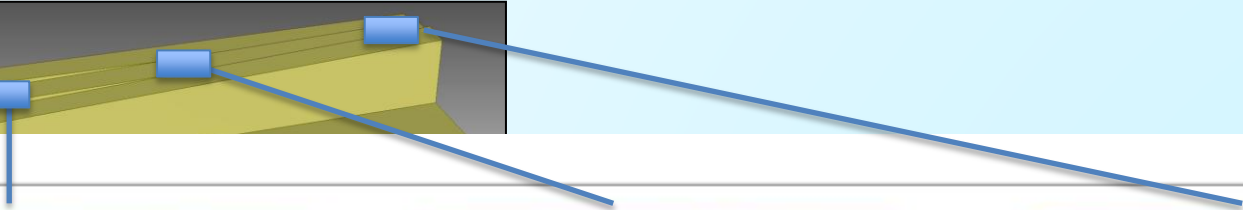
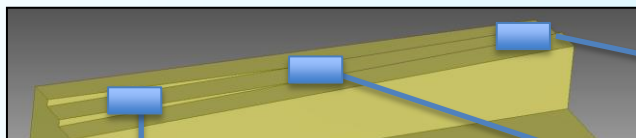
- difficult to make a non-spherical surface
- cause the optical axis deviation under low-T

# Test fabrication

- ✓ All mirror in SWIMS-IFU will be made of aluminum alloy with high precision machining technique



match the thermal expansion with support structures under a cryogenic environment.



# Present condition of multi-object IFU

✓ KMOS-VLT (first light occurred on 21 November 2012)



- $\lambda=0.8 - 2.5 \text{ um}$   
( $R=3000-4000$ )
- patrol field:  $\Phi 7.2'$
- 24 IFUs  
(each FOV is  $2.4'' \times 2.4''$ )
- spatial sampling:  $0.2'' \times 0.2''$

They adopted **the advanced image slicer IFUs.**

# Future instruments using IFS

## <Ground-based observation>

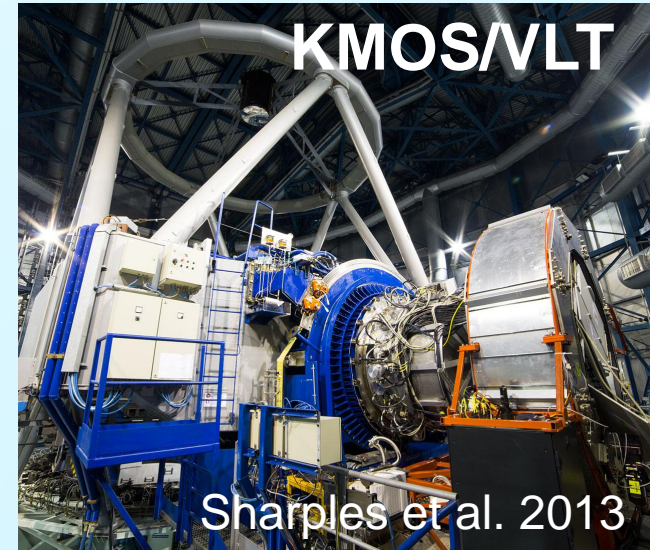
✓ VLT

- KMOS(near-IR), MUSE(optical)

✓ TMT

- IRIS(near-IR),

They have also lenslets mode.



## <Space observation>

✓ JWST

- NIRSpec(near-IR)

**The advanced image slicer IFUs  
is the key technique for future instruments**



## Questions to Speakers

To all workshop speakers: please include your responses to the questions below.

### Questions on instrument specifications

#### Primary questions

- (Q.1) In the baseline specifications of NIR instruments, what are the most important instrument options?
  1. Wide-Field Near-IR Imager
  2. Wide-Field NIR Imager and Multi-Object Spectrograph
  3. Multi-Object Integral Field Spectrograph

Which instrument is *essentially* important for your science cases?

I'd like to propose  
the 4<sup>th</sup> option !

“**MOS camera + IFU**”

## Wide-field NIR imager & MOS with GLAO + IFS mode

- ✓ This type of IFU (e.g. SWIMS-IFU, GNIRS, v-MOIRCS?) is a fully independent , self-contained module

There are no IFUs with GLAO for now,  
but above will give the way to achieve it.

# Summary



- ✓ **Integral Field Spectroscopy** have an increasing impact on our understanding of galaxy evolution mechanisms.
- ✓ **The advanced image slicer IFUs is the key technique** for future instruments such as TMT, JWST (Of course for Subaru !!)
- ✓ **SWIMS-IFU is now under developing**
  - constructing the solution which satisfies the size limit  
In this case, slice mirror has curvature
  - verifying the high precision machining technique by fabricating the test pieces